

Modeling the Production, Delivery, and Maintenance of OH/H₂O on the Lunar Surface

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Acknowledgments:

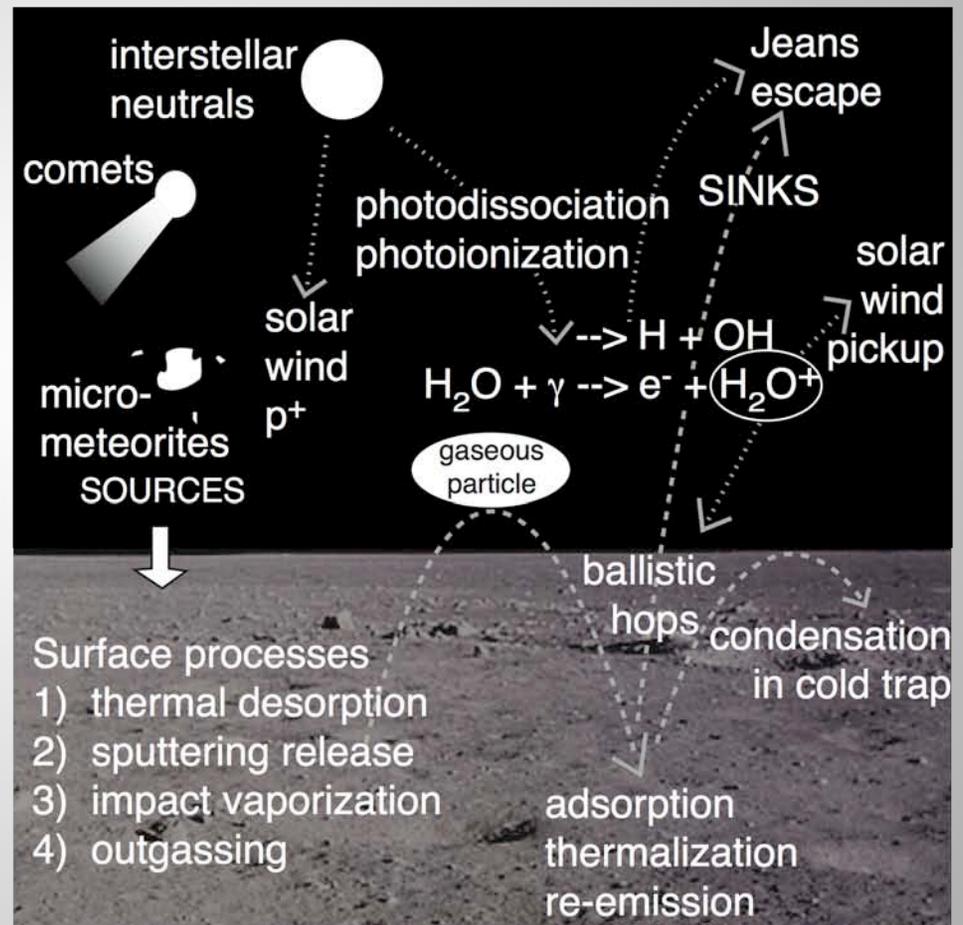
NASA Lunar Science Institute, NASA
PGG and NASA LASER

The logo for Applied Physics Laboratory (APL) consists of the letters 'APL' in a large, bold, sans-serif font. The letters are dark red and are positioned on the right side of the slide, above the university name. The background of the slide features a faint grid pattern in the upper right and a technical drawing of a lunar lander in the lower left.

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APPLIED PHYSICS LABORATORY

Lunar Surface-Atmosphere System

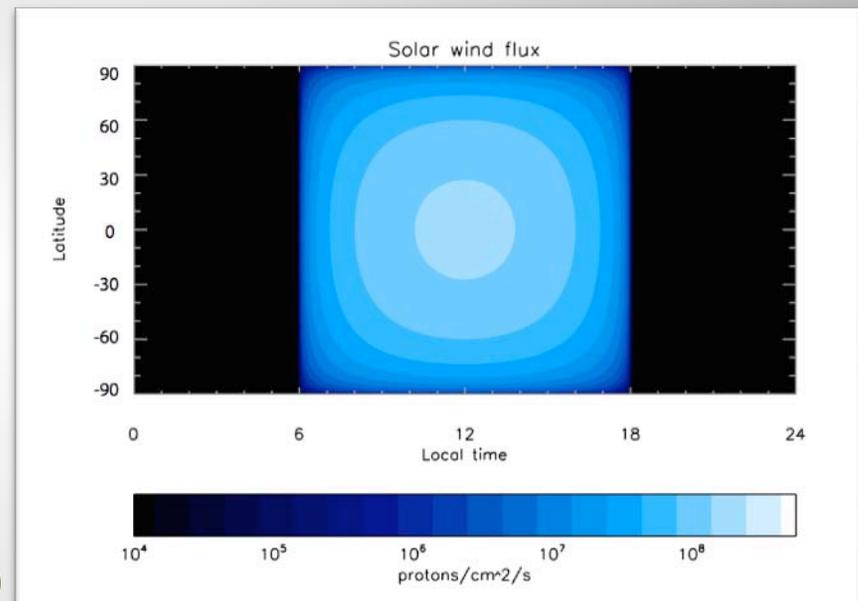
- **Source rate**
 - Dust, comets, solar wind
 - Distribution and velocity
- **Surface stability**
 - Thermal desorption
 - Photon-stimulated desorption
 - Sputtering
 - Impact vaporization
- **Migration**
 - Residence time
 - Lifetime
 - Surface redistribution
 - Inward migration
- **Loss mechanisms**
 - Photodissociation
 - Sputtering
 - Surface sinks



Background photo from N.S.S.D.C., Apollo 11 landing site

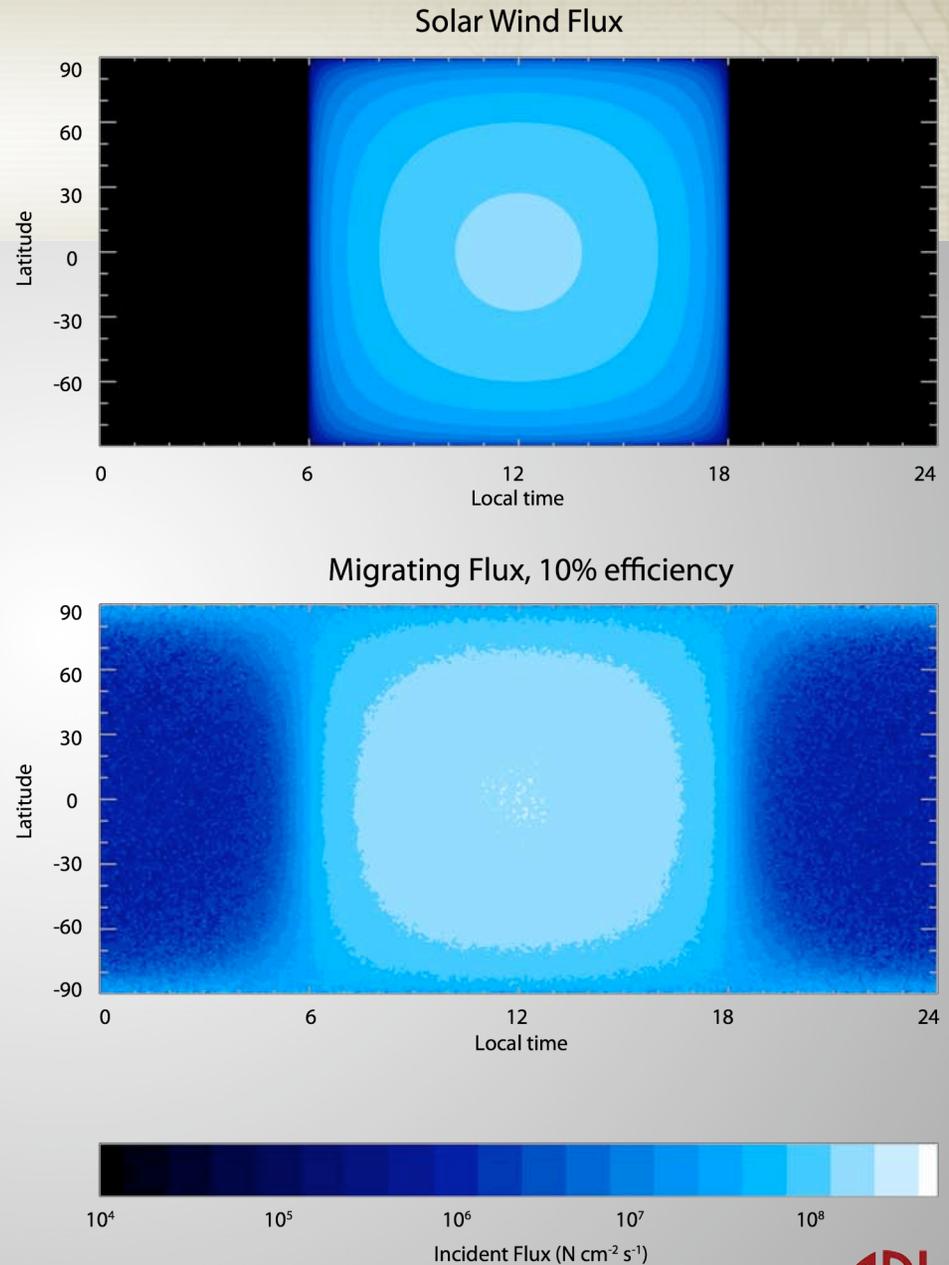
Solar Wind (SW) Delivery Rate

- SW flux falls off with solar zenith angle (SZA).
- The distribution does not match observations of high amounts at the poles.
- Converting 100% SW to H₂O for 1 lifetime:
 - Fluence of $3.4e12$ H₂O/cm²
 - Convert to global layer 10^{-10} cm.
- Assuming a weight fraction of 1000 ppm, and 100% conversion to H₂O, a thickness of 10 Å can be produced in steady state.



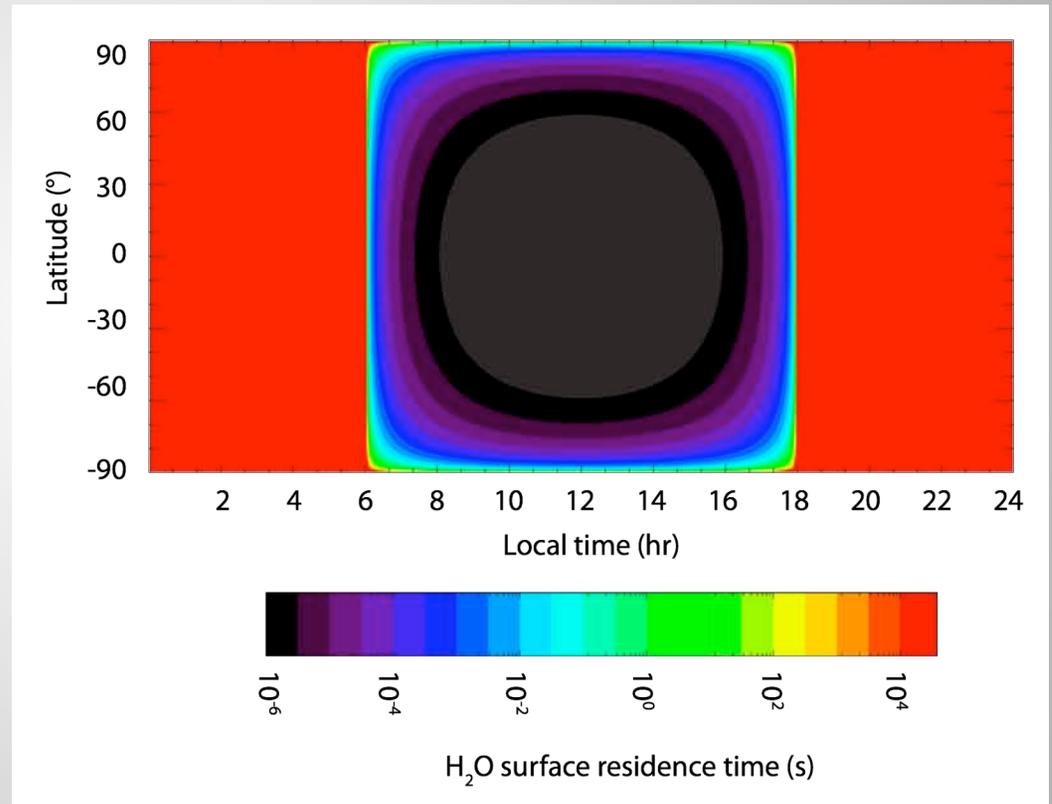
Migrating Flux

- Monte Carlo model of migration pattern of H₂O
- Each particle takes an average of 27 hops—boosts flux
- Migration redistributes particles to higher solar zenith angles
- Still have the peak flux at the equator



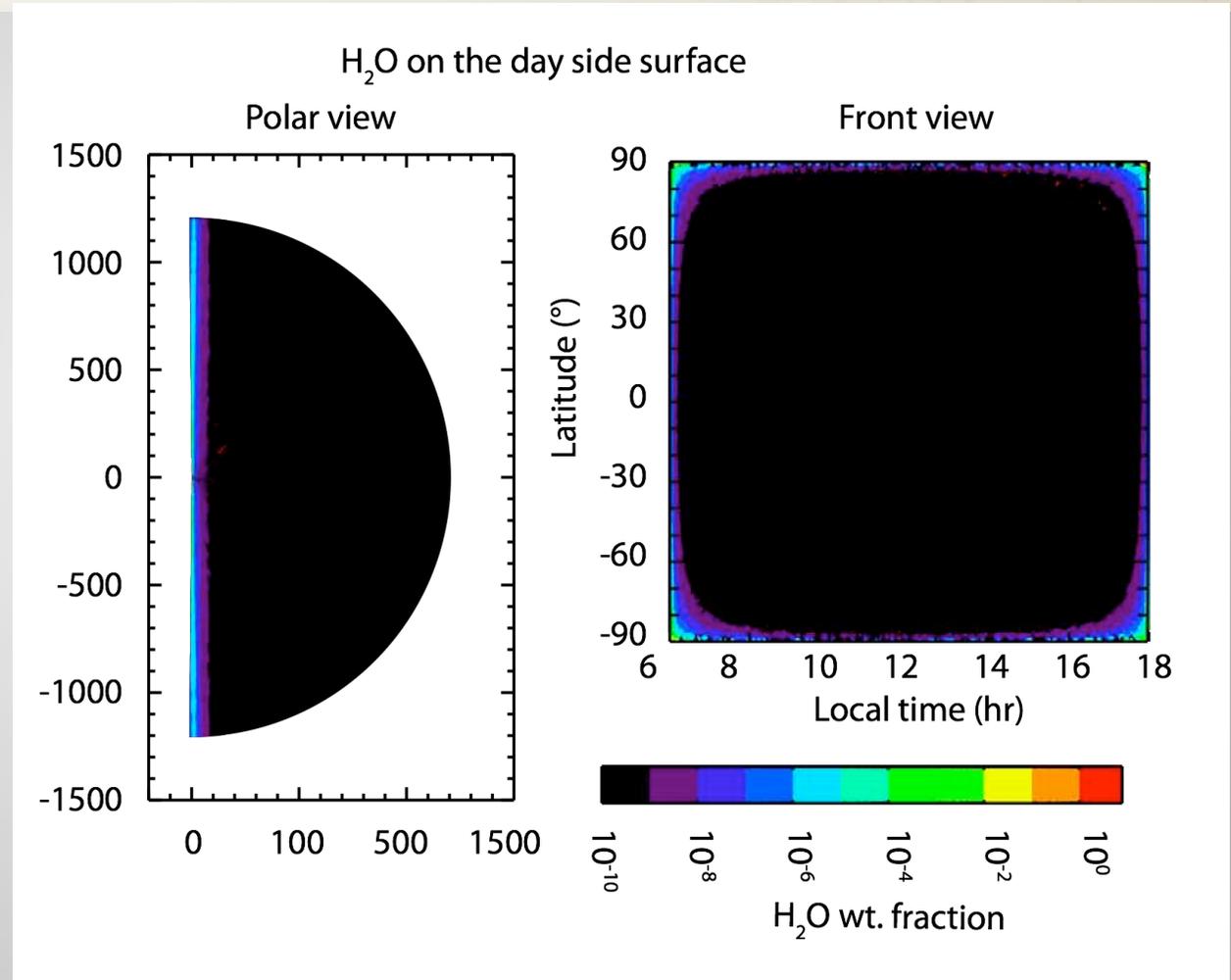
Surface Residence

- Migrating H₂O particles reencounter the surface, interact, and later thermally desorb
- Observations are of particles on the surface
- Short dayside residence times, long nightside residence time



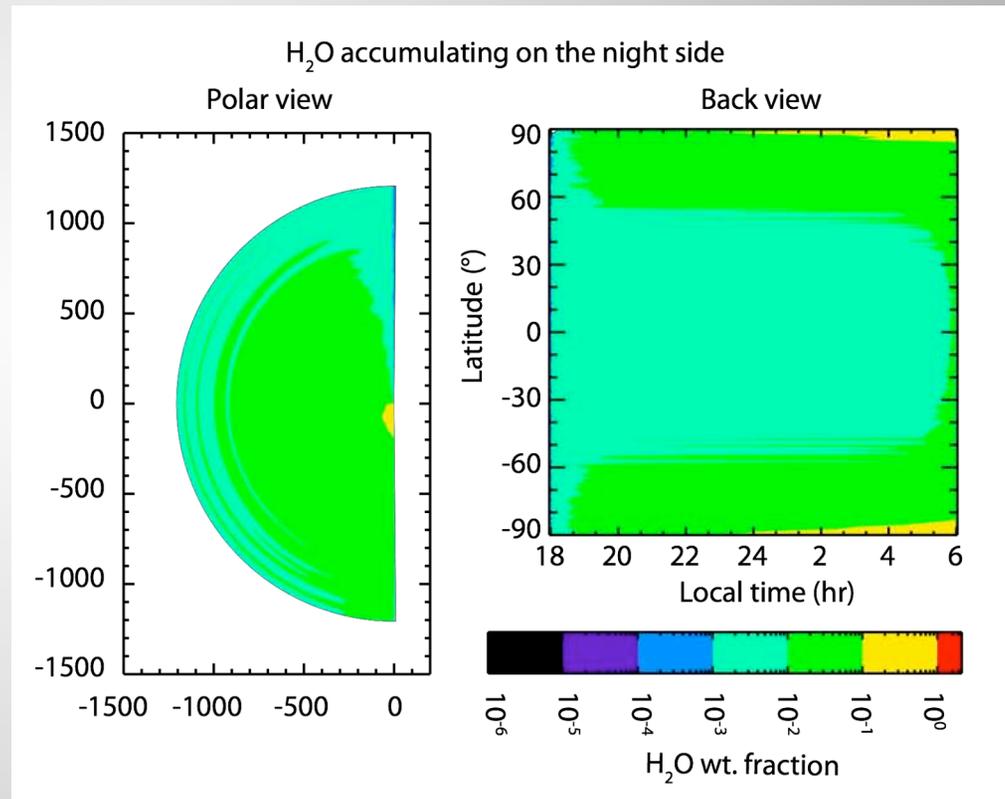
Dayside H₂O Results

- Convolve migrating flux with a surface residence time function
- Result is a distribution that has higher concentrations at high latitude
- Weight fraction predicted is too low to observe except very close to the terminator ($1e^{-3}$).



Adsorption to the nightside

- Long residence time on nightside due to low UV flux and low temperature
- Depending on release conditions, ~40% of molecules get trapped on the nightside each lunation.
 - Average particle lasts 1.67 cycles
- Migration on grain surfaces into regolith is possible
 - Perhaps this sequesters H₂O released by solar wind on the last lunation to depths below UV penetration



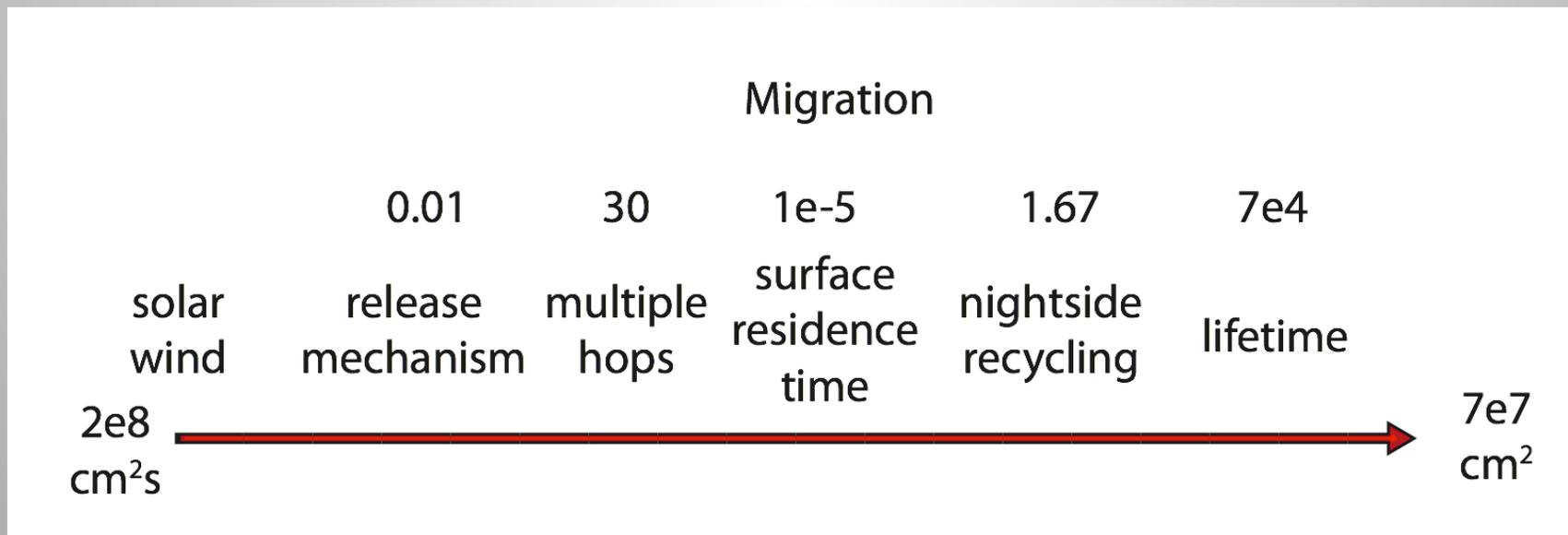


Conclusion

- Solar wind in steady state does not reproduce observed H₂O surface density
- Micrometeoroids—source rate of $2e5 \text{ H}_2\text{Ocm}^{-2}\text{s}^{-1}$
 - Did a simulation with a more global release distribution—changes sink on nightside and surface density by 10-25% from simulations shown here
- Comets—source rate of $3e6 \text{ H}_2\text{Ocm}^{-2}\text{s}^{-1}$
 - Sporadic
- Wild speculation
 - Recent larger source (comet)
 - Seasonal sequestration
 - Thick exchange layer perpetuated by diurnal cycle
 - Complex surface chemistry/long residence times

Future Directions

- Surface atmosphere interactions
- Regolith reservoir/depth recycling
- Stochastic sources

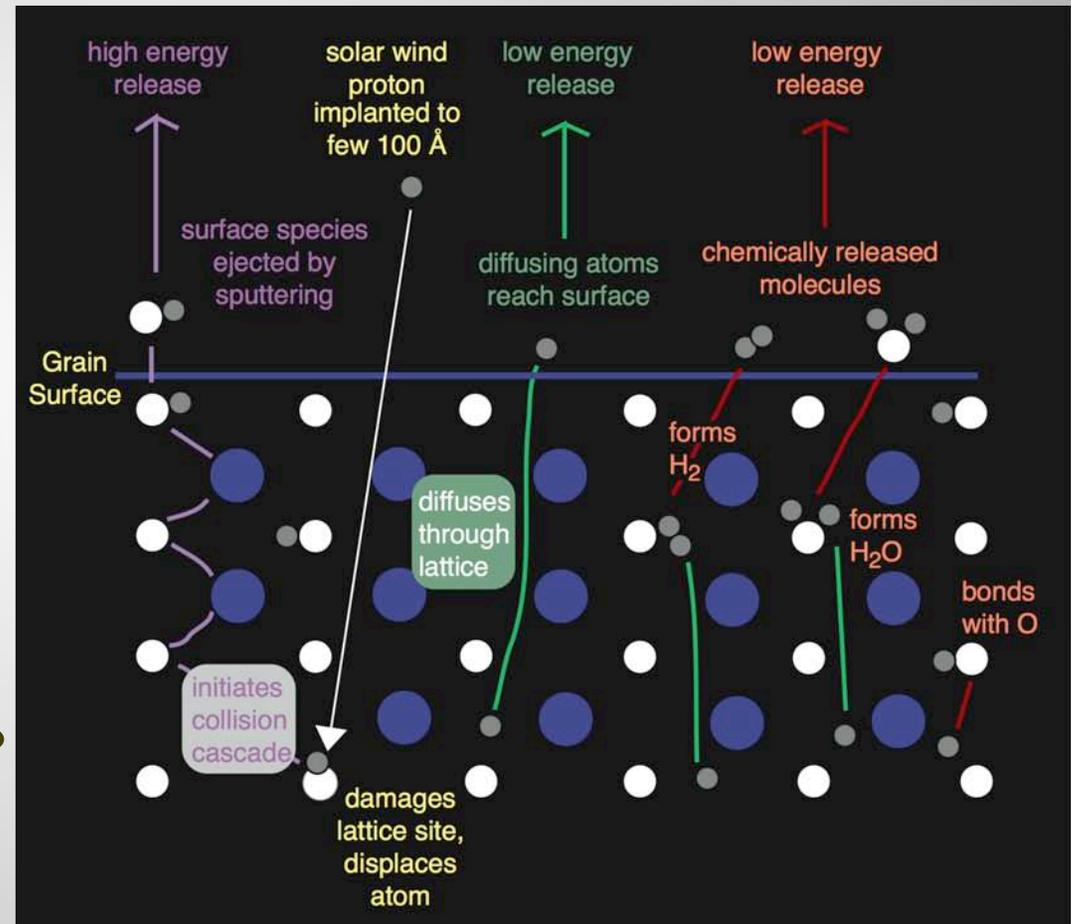




Backup

Solar Wind Implantation

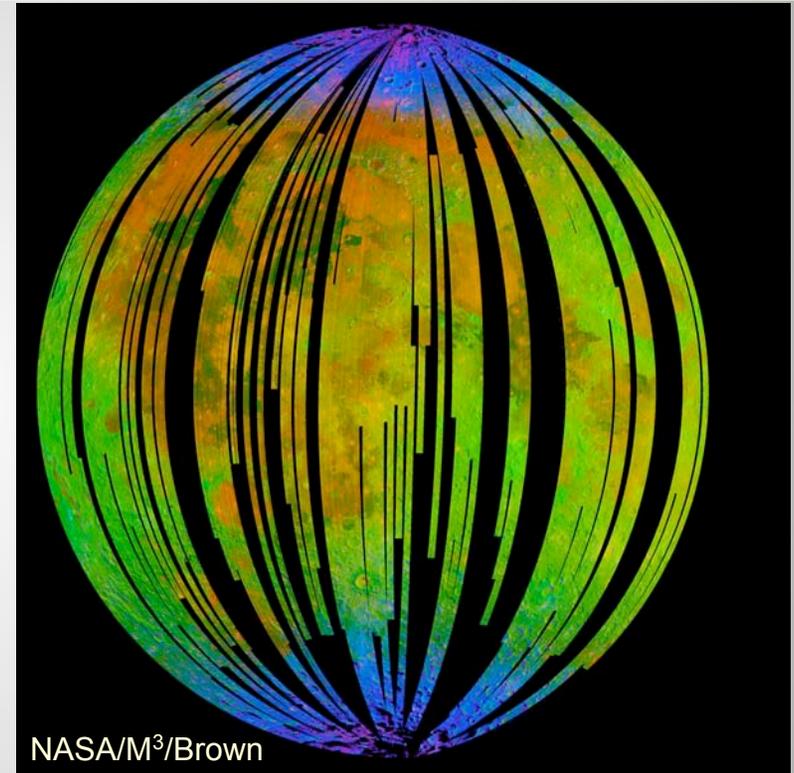
- Solar wind protons are one possible source of hydrogen for OH/H₂O
 - Protons penetrate ~200 Å
 - Radiation damages lattice
 - Hydroxyl formation
- Lab indications are that saturation levels achieved in fluences expected in much less than the regolith turnover time
 - Mature regolith saturated?
- Perhaps there are two components, a physisorbed and a chemisorbed.



OH/H₂O Observations

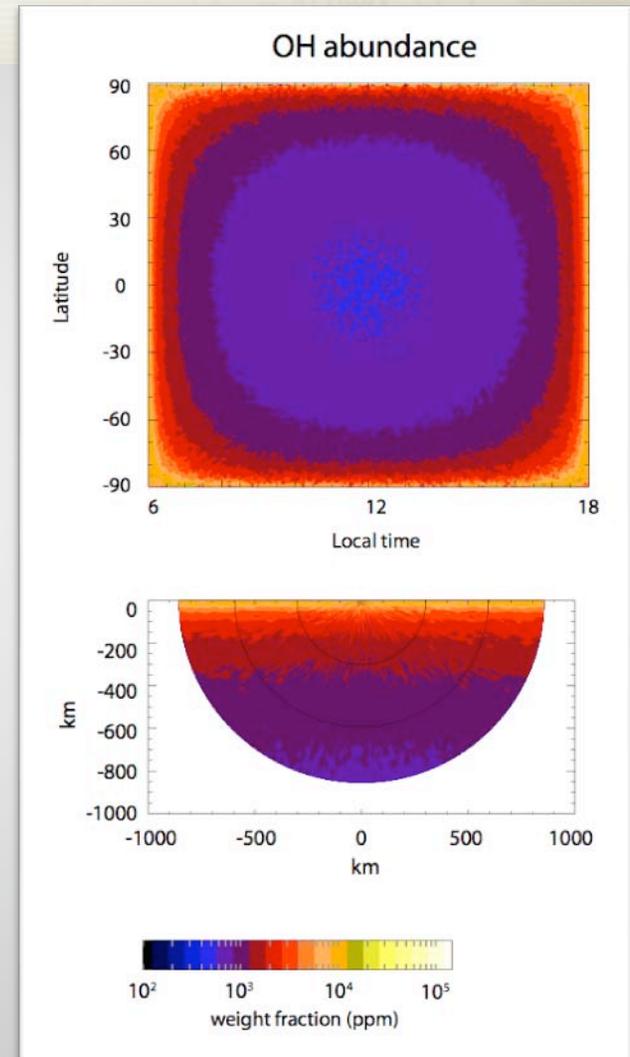
Groundbreaking new observations bring new questions

- **Contents**
 - Is it OH, H₂O, or both?
 - How much?
- **Distribution**
 - Lunar compositional dependence?
 - Solar zenith angle dependence?
 - Depth?
- **Physics**
 - Source?
 - Maintenance?
 - Loss?

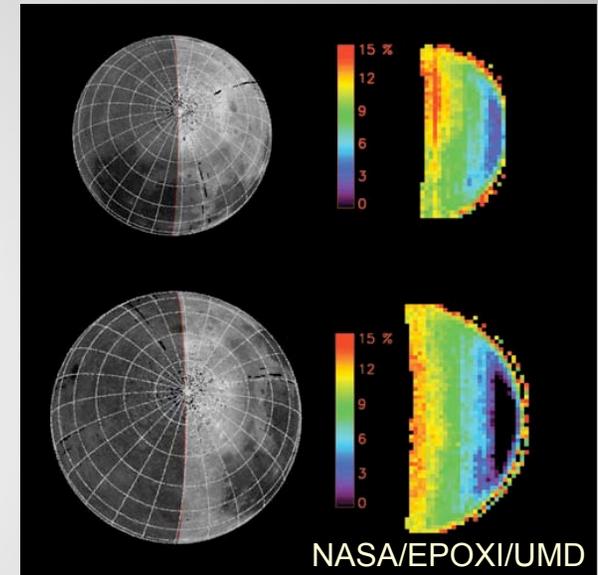
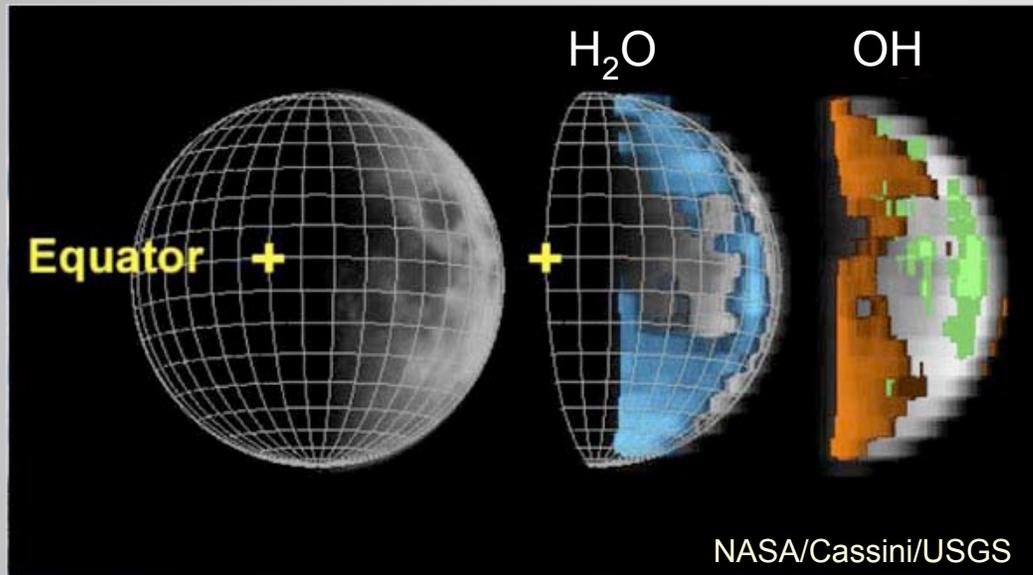


OH Simulations

- This simulation reproduces a diurnal surface density when pushed to these limits:
 - 100% efficiency of SW conversion to OH/H₂O
 - 500Å layer depth
 - Subsolar residence time of 18 hrs
 - Scales to a monolayer on 10 μm grain of 30 ppm



Spatial Distribution Lunar OH/H₂O



- Highest concentrations are at highest latitudes (all three)
- EPOXI data indicate a diurnal signal (Sunshine et al., 2009)
- Investigate the possibilities of a diurnal signal
 - modeling the surface-atmosphere system
 - Start simple, add pieces one at a time
 - (Crider and Vondrak, 2000; 2002)