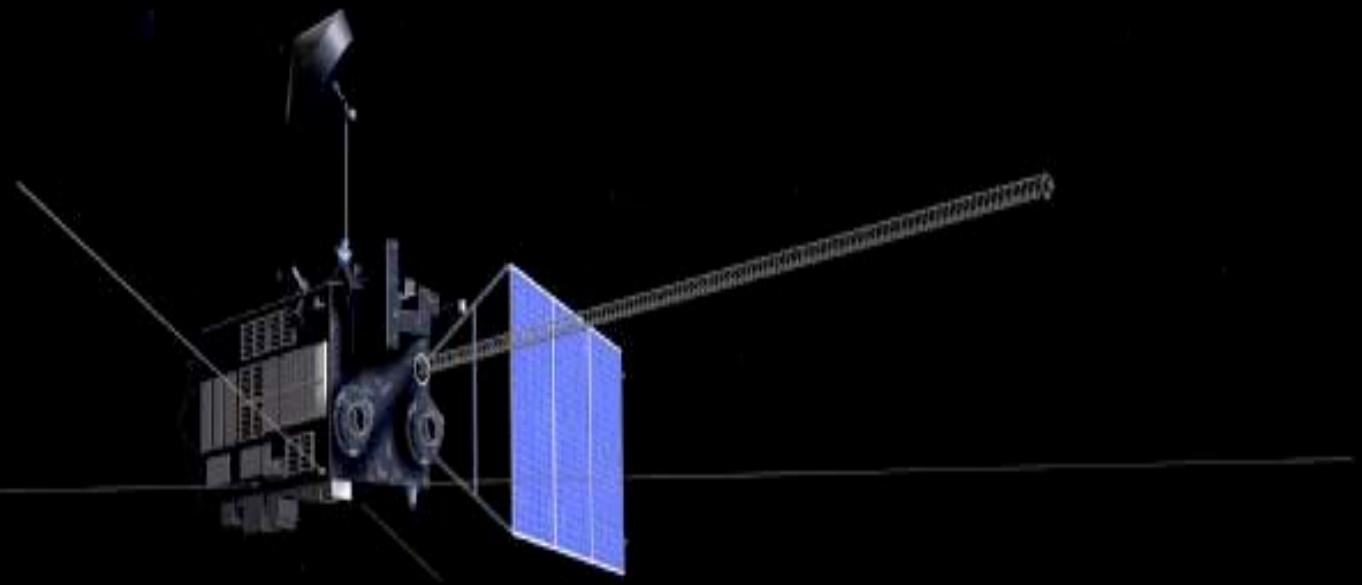




Comparison of PKT and SPA Regions of the Moon Revealed Through KGUYA GRS



Kyeong Ja Kim & co-authors

Korea Institute of Geoscience and Mineral Resources, Korea.



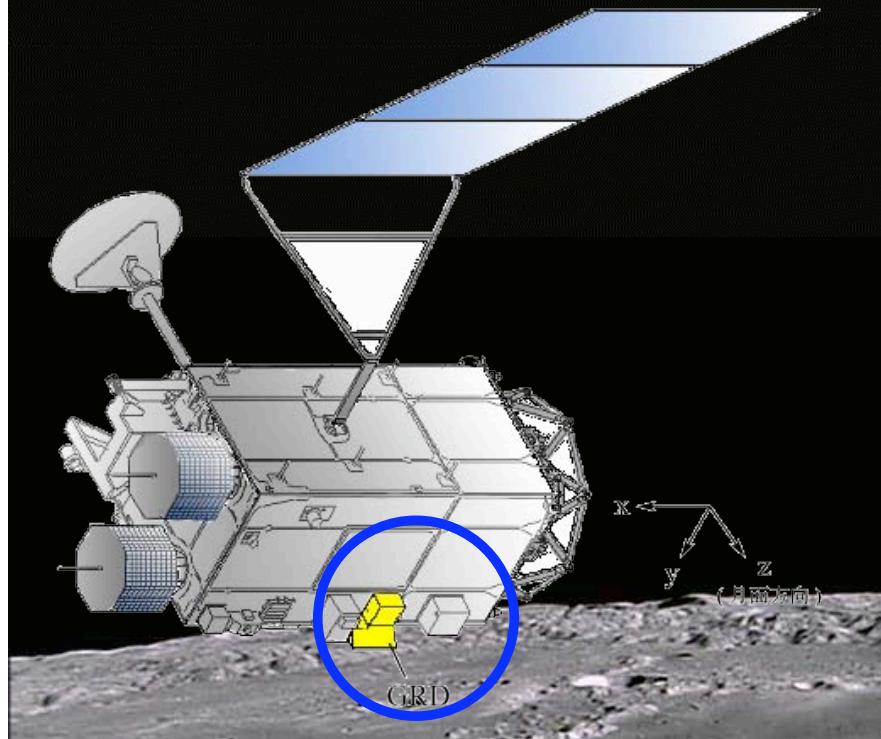
**K. J. Kim¹, J. M. Dohm², J.-P. Williams³, J. Ruiz⁴, B.-H. Yu⁵,
T. M. Hare⁶, N. Hasebe⁷, N. Yamashita⁷, Y. Karouji⁷, S.
Kobayashi⁸, M. Hareyama⁸, E. Shibamura⁹, M. Kobayashi¹⁰,
C. d'Uston¹¹, O. Gasnault¹¹, O. Forni¹¹, Reedy¹²**

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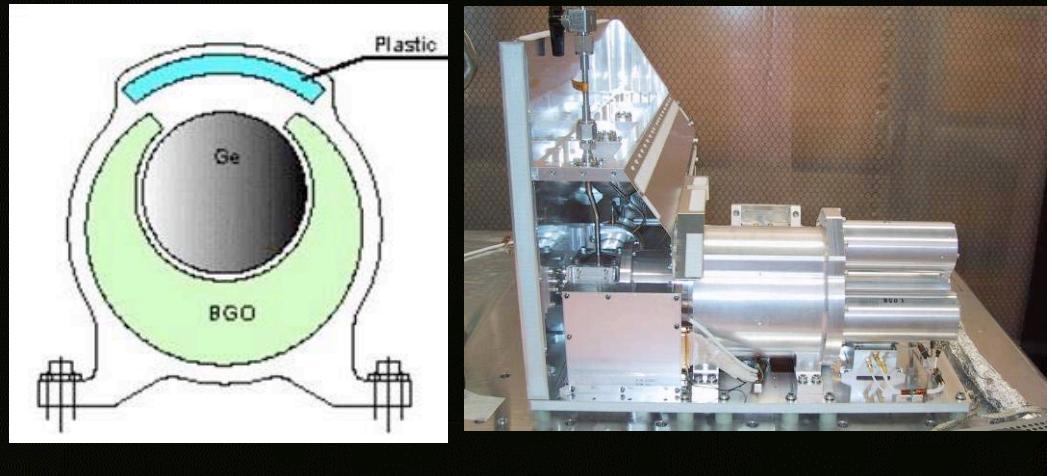


Kaguya Gamma-Ray Spectrometer

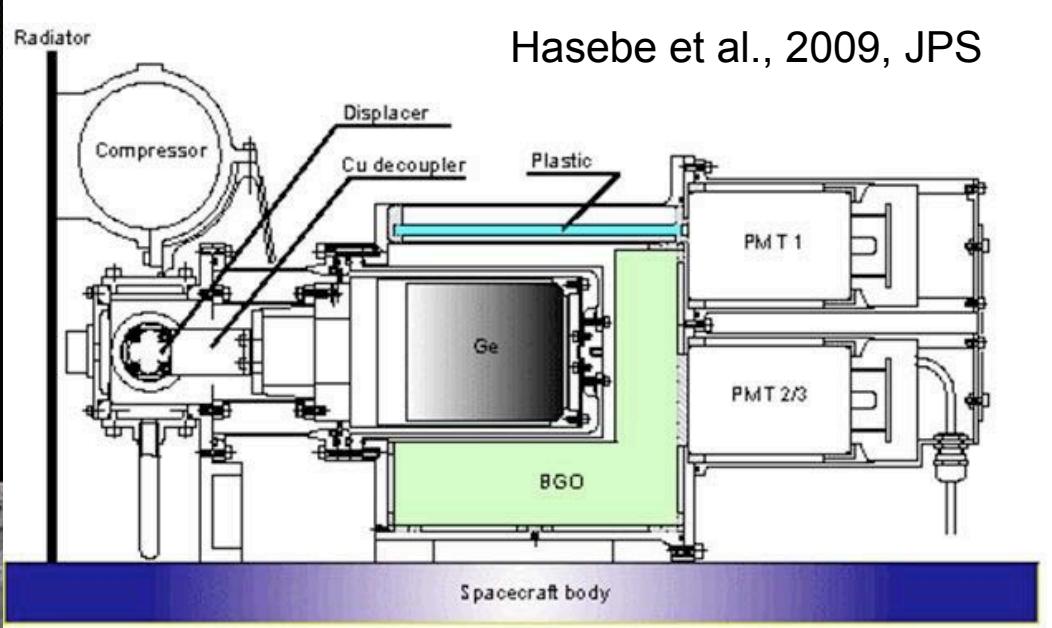
- ❖ Main Detector
 - HP-Ge(Eurisys Measures)
- ❖ Anti-Coincidence Detector
 - Plastic Scintillator(BICRON)
 - BGO (BICRON)
- ❖ Global mapping of K, U, Th, O, Mg, Al, Si, Ca, Ti, Fe, and H



KAGUYA GRS with high energy resolution



Hasebe et al., 2009, JPS





Comparison with missions in the past

Mission	Launch	Main detector		Anticoincidence detector	Mount
		Material	Size [mm]		
Apollo 15/16	1971/1972	NaI(Tl)	φ76×76	PS	7.6m boom
Lunar Prospector	1998	BGO	φ71×76	PS	2.5m boom
Kaguya	2007	Ge	Φ65 × 77	BGO+PS	on SC body

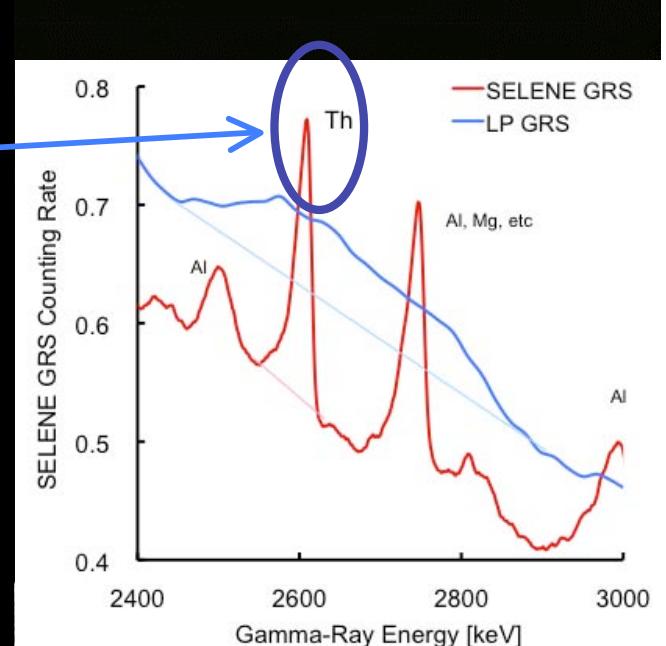
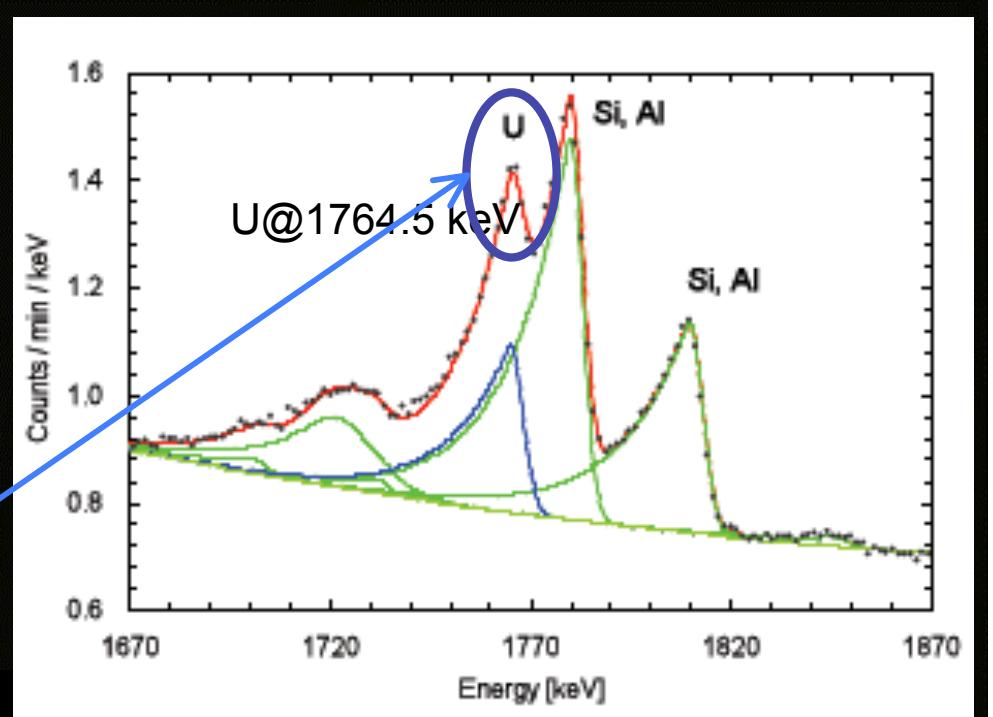
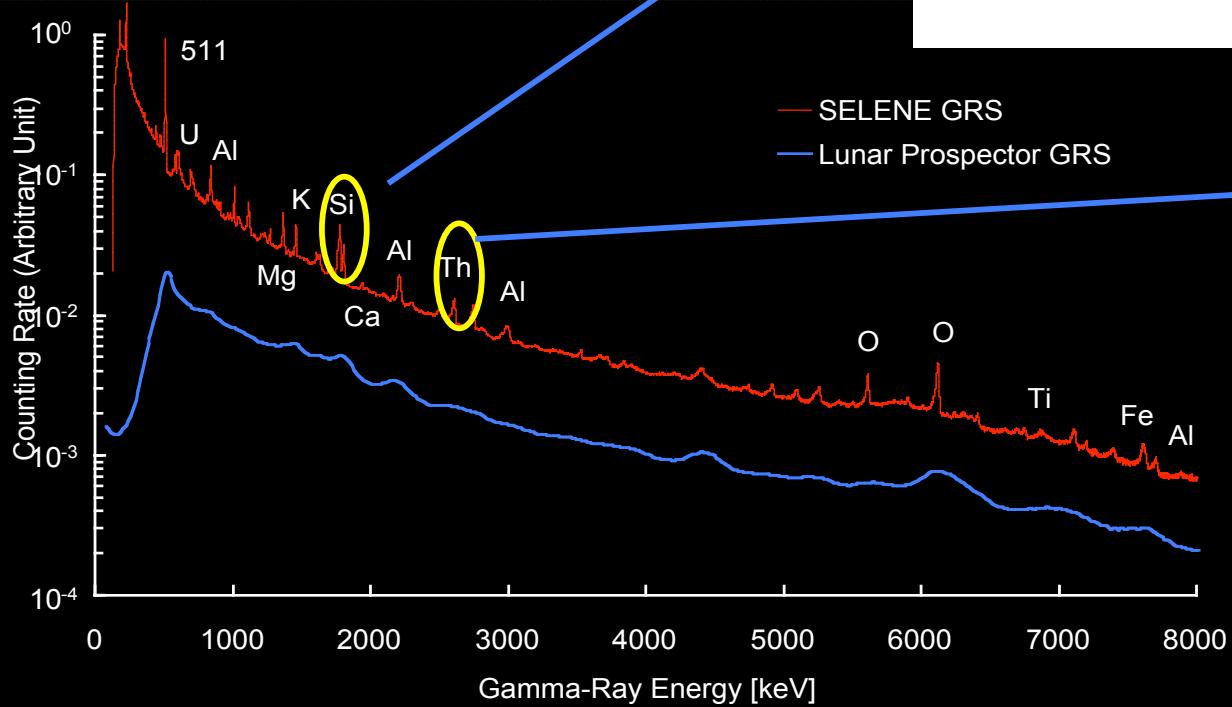
	NaI(Tl) φ3" ×3"	BGO φ3" ×3"	Ge 250 cm ³
Density (g/cm ³)	3.67	7.13	5.33
Efficiency*	1	1.8	0.6
FWHM@1.33MeV [keV]	46	74	3
Figure of merits*	1	1.4	4.2

normalized to NaI(Tl)



Evidence of Excellent Identification

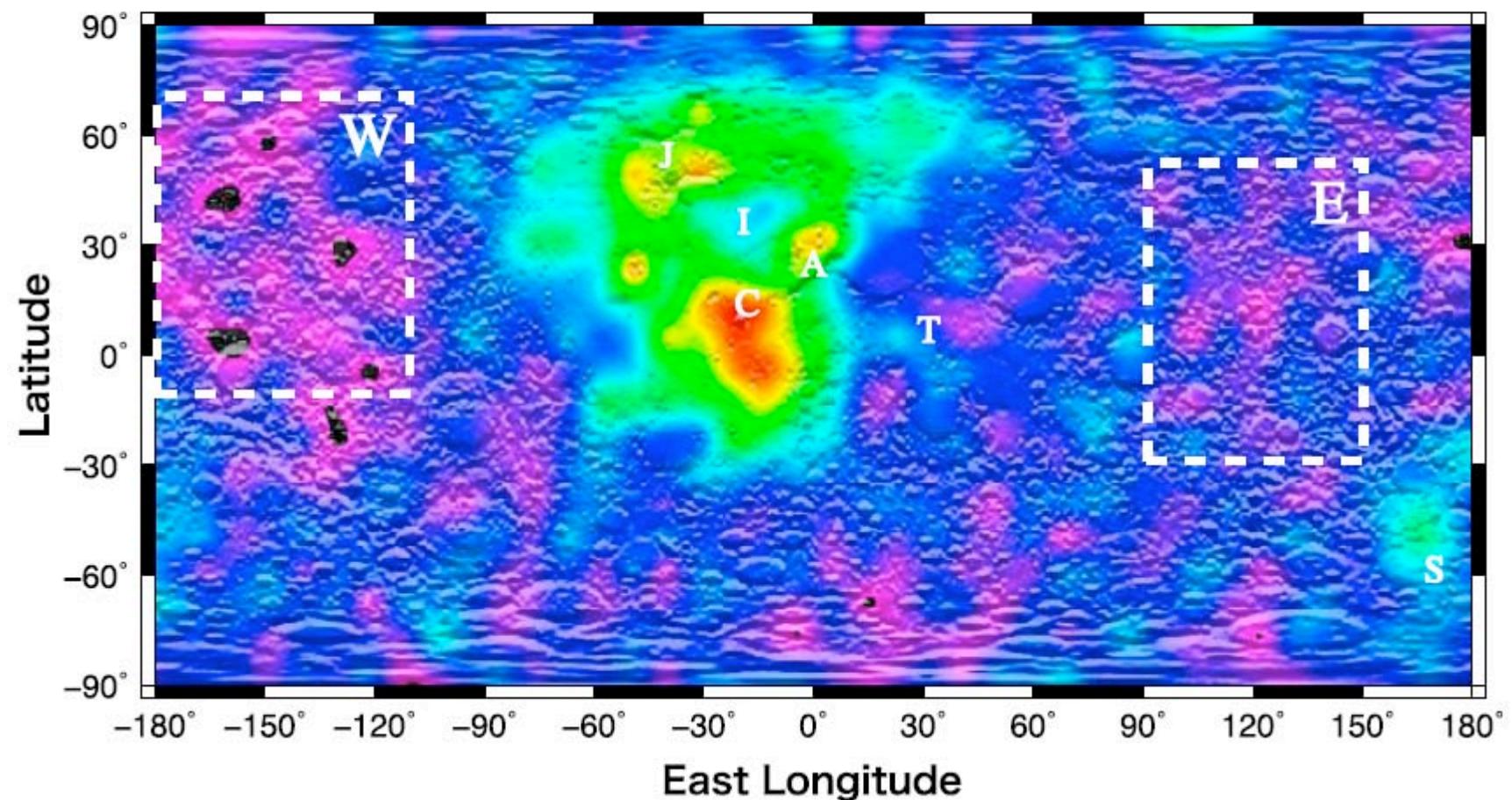
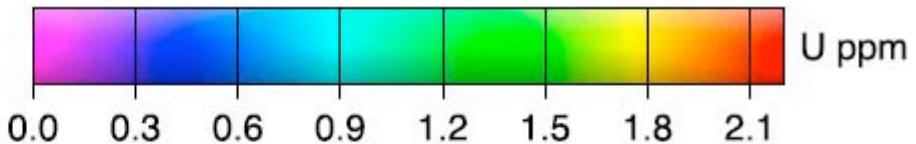
Fitting analyses of gamma-ray lines
in the energy spectra of KGRS from
U at 1765 keV and
Th from 2615 keV.





U Map by Kaguya Gamma-Ray Spectrometer

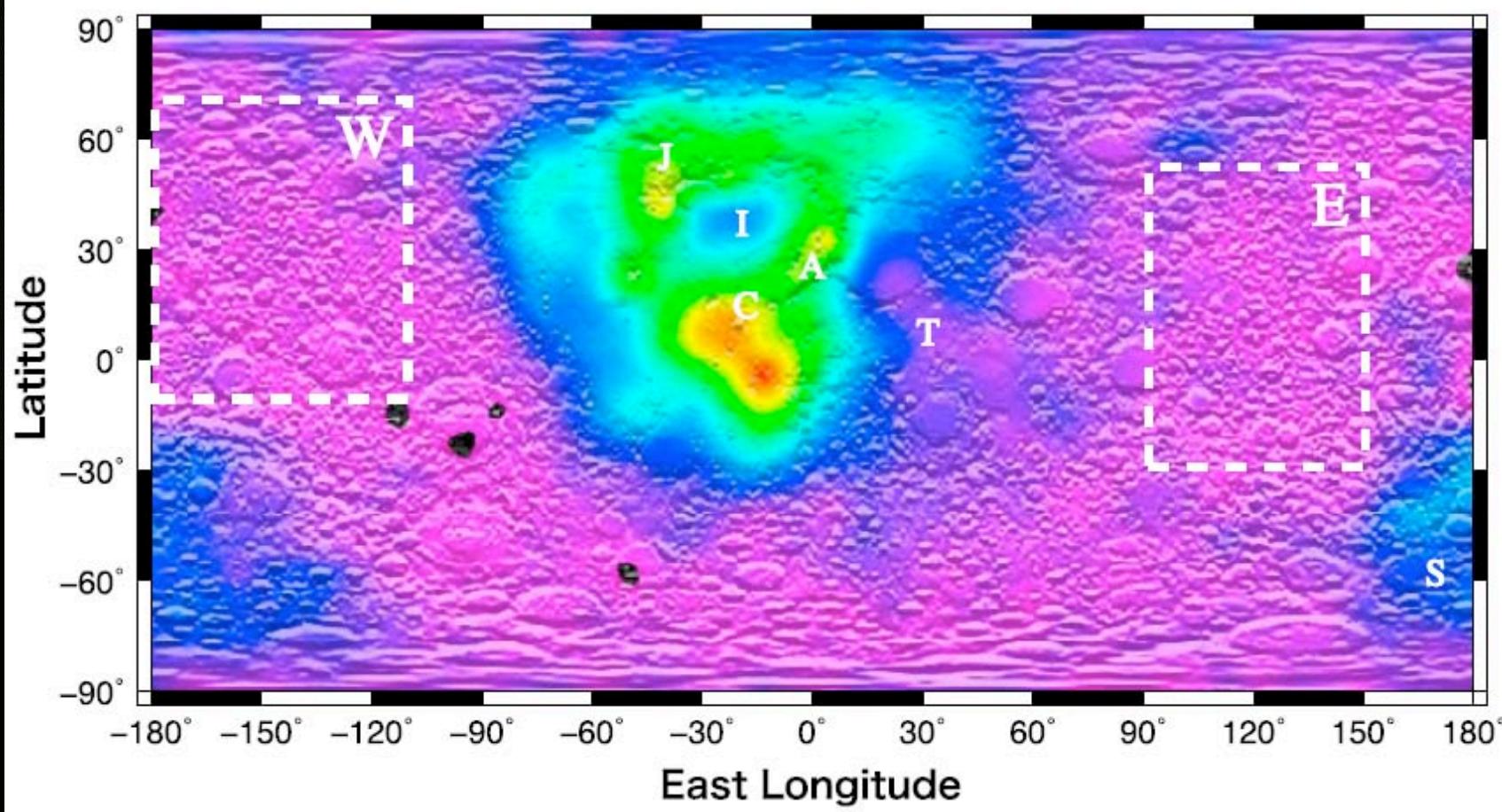
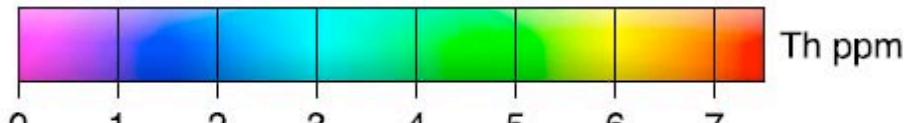
N. Yamashita et al., 2010 (GRL)





Th Map by Kaguya Gamma-Ray Spectrometer

N. Yamashita et al., 2010 (GRL)





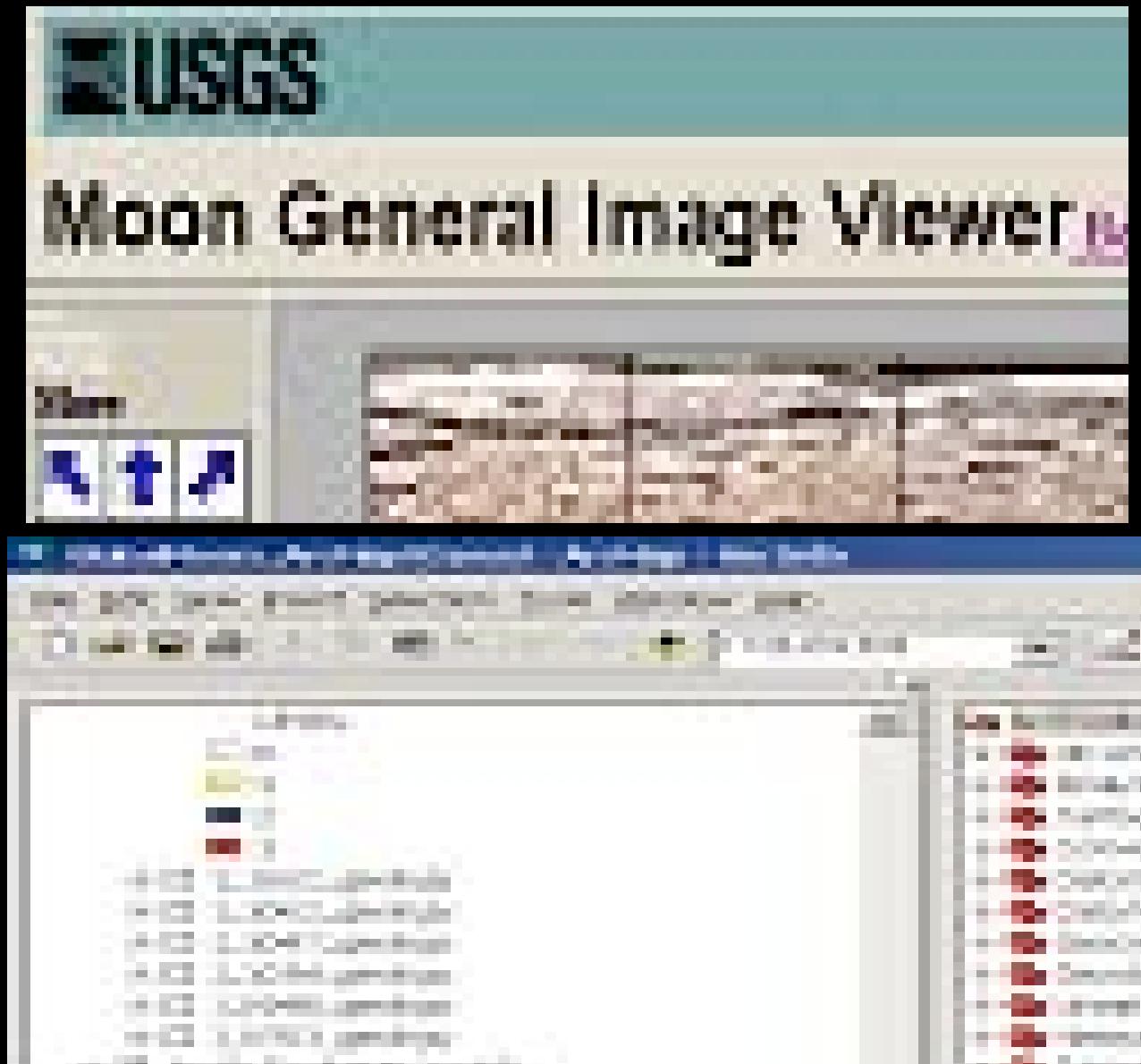
Methodology

Comparison among Procellarum and South Pole Aitken Basin Regions

< Comparative analysis among lunar geology and Kaguya elemental information >

- 1. Define regions of scientific interest
- 2. Compile of Apollo-based stratigraphic information of sites
- 3. Compile elemental information of sites of the scientific interest
- 4. Perform comparative analysis among the stratigraphic and elemental information (i.e. cookie cut out)
- 5. Interpret surface evolution through time

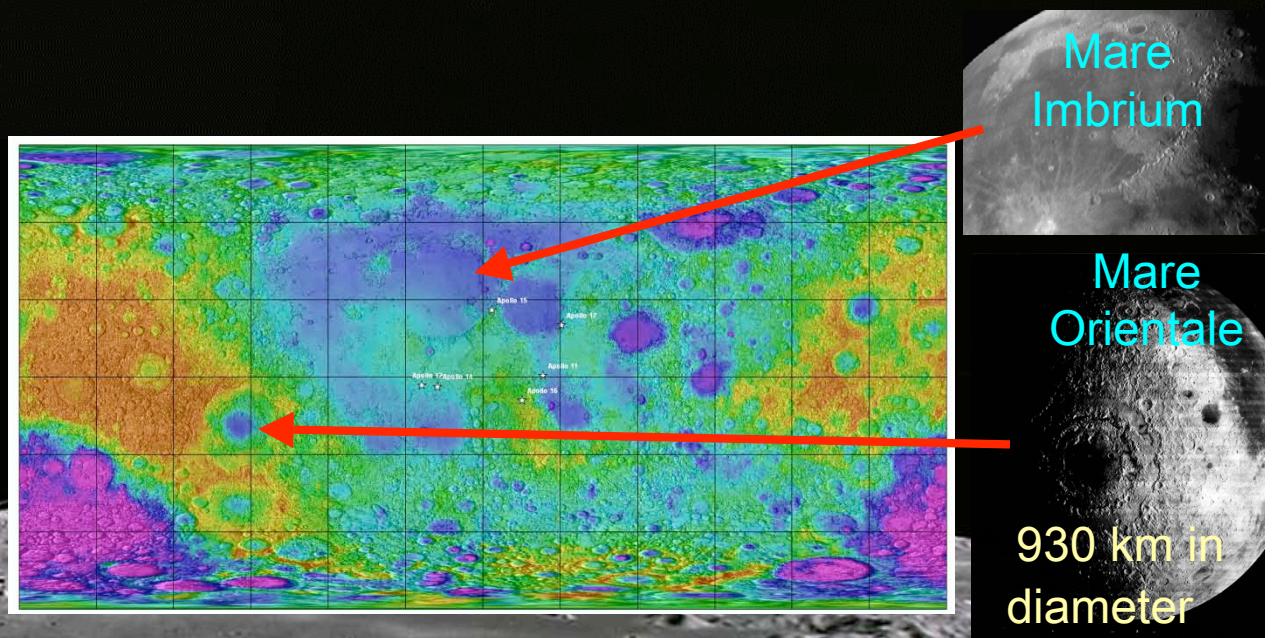
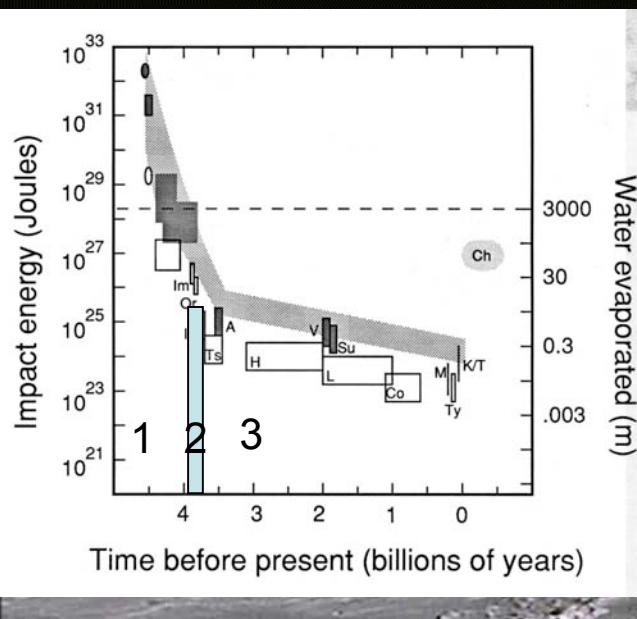
An ArcView lunar GIS mapping system introduced. Planetary GIS-Interactive Map Analysis Program: PGIMAP





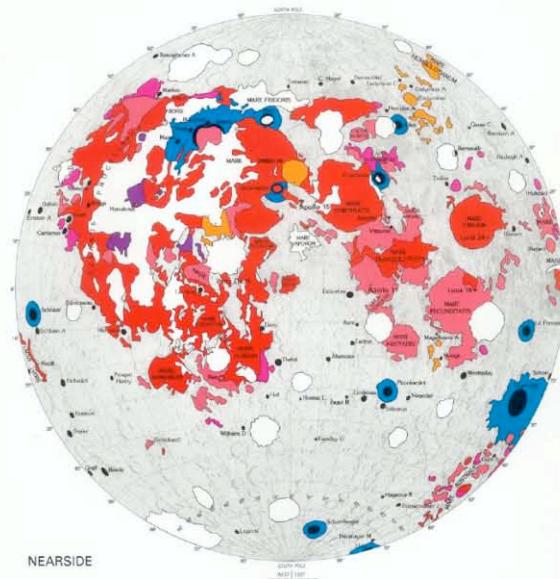
Age information of the Moon

Stage	Age	Period
1	4.5 - ~4.0 GA	Period of heavy bombardment
2	~3.9 - ~3.7 GA	Period of Late Heavy Bombardment (~3.9 - ~3.7 GA), including possible period of impact catastrophism
3	~ <3.7 GA	Post late heavy bombardment, mare volcanism (~ <3.7 GA)



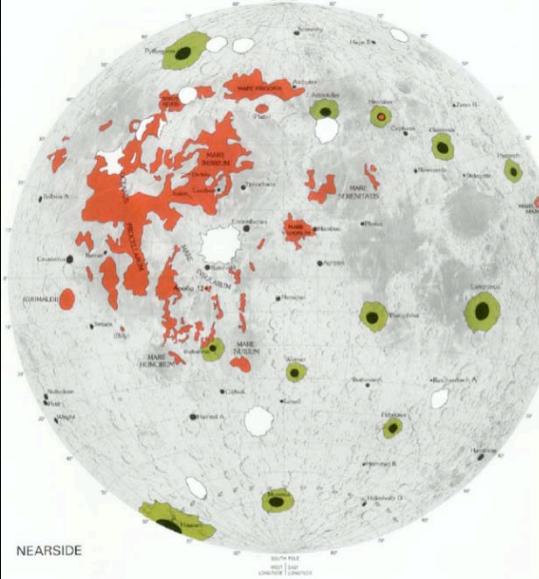


**Stage 3:
Upper Imbrian System**



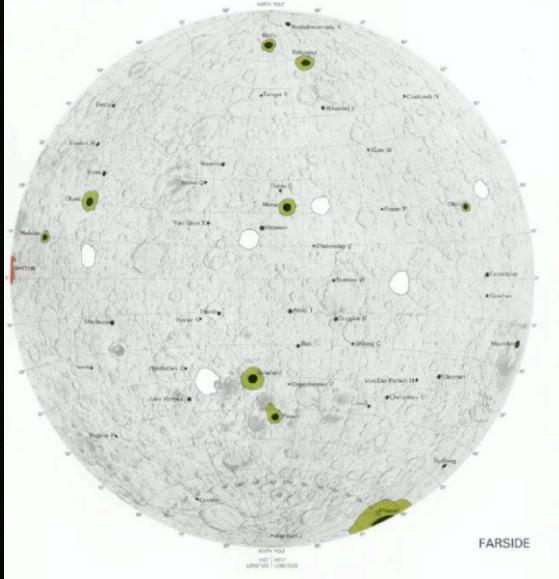
NEARSIDE

**Stage 3:
Eratosthenian System**

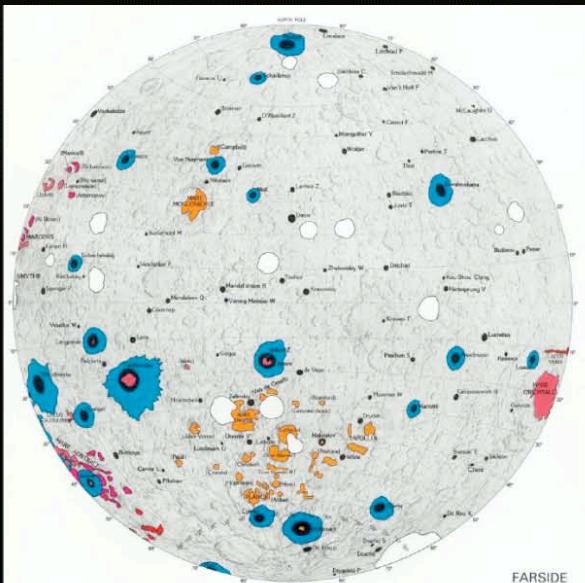


NEARSIDE

**Stage 3:
Copernican System**

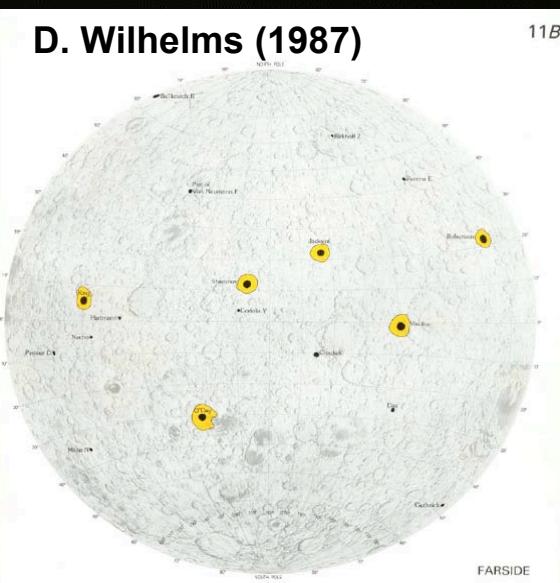


FARSIDE

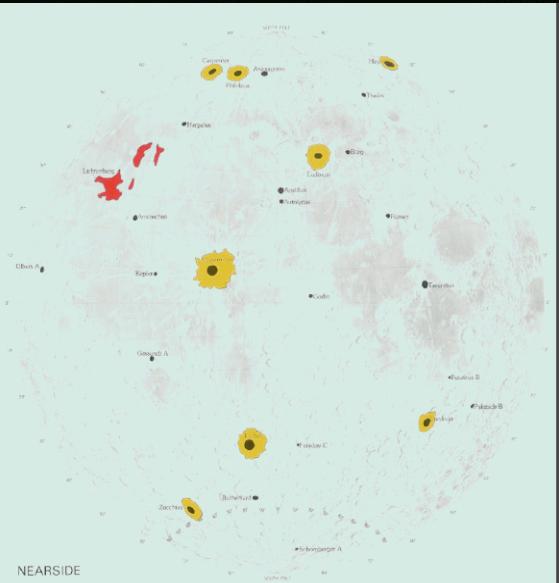


FARSIDE

D. Wilhelms (1987)



11B

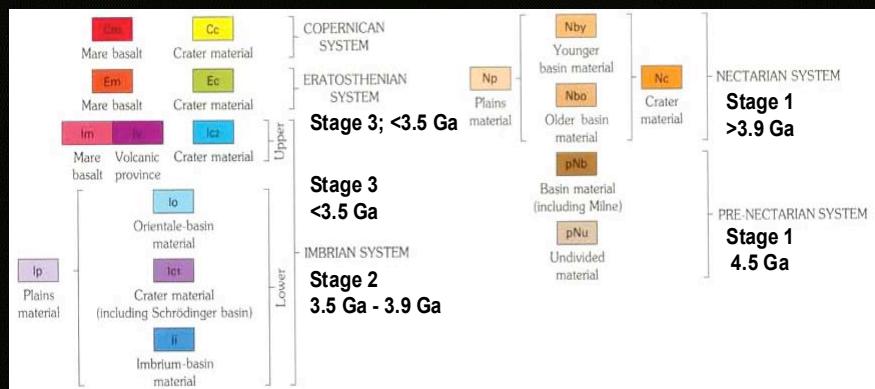
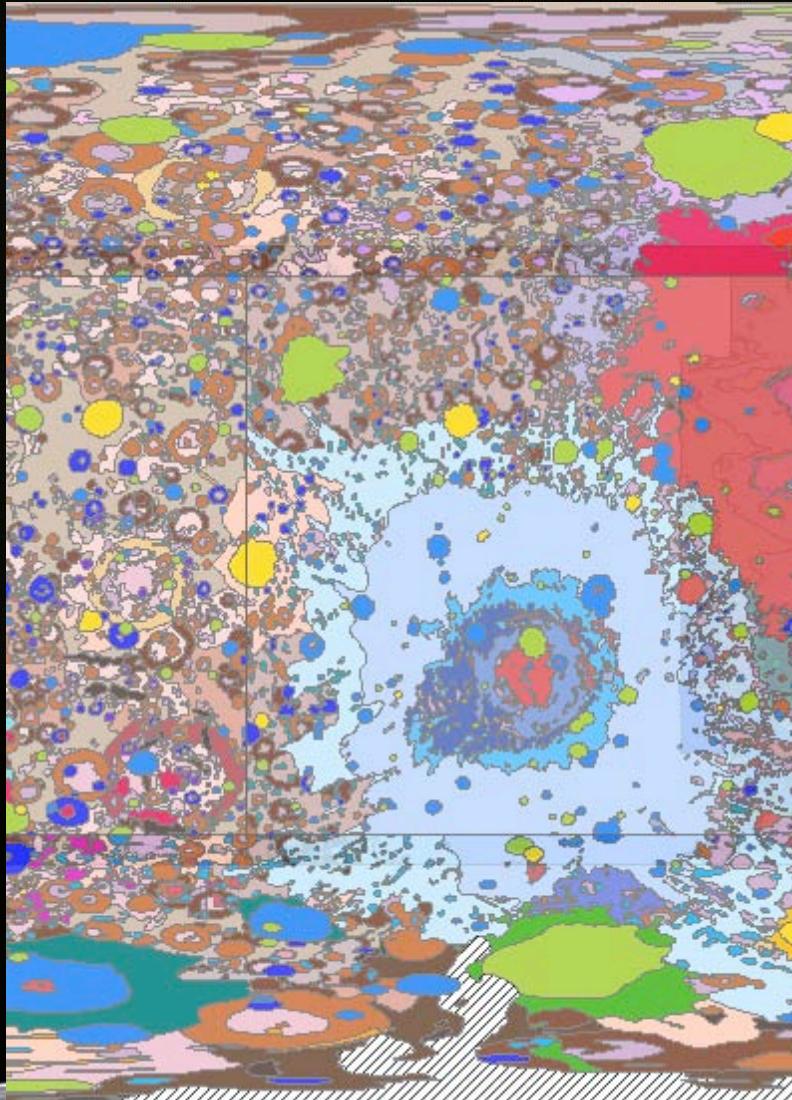


NEARSIDE



GIS based geologic map

Wilhelms and McCauley, 1971; Wilhelms and El-Baz, 1977; Scott et al., 1977; Stuart-Alexander, 1978; Lucchitta, 1978; Wilhelms et al., 1979.

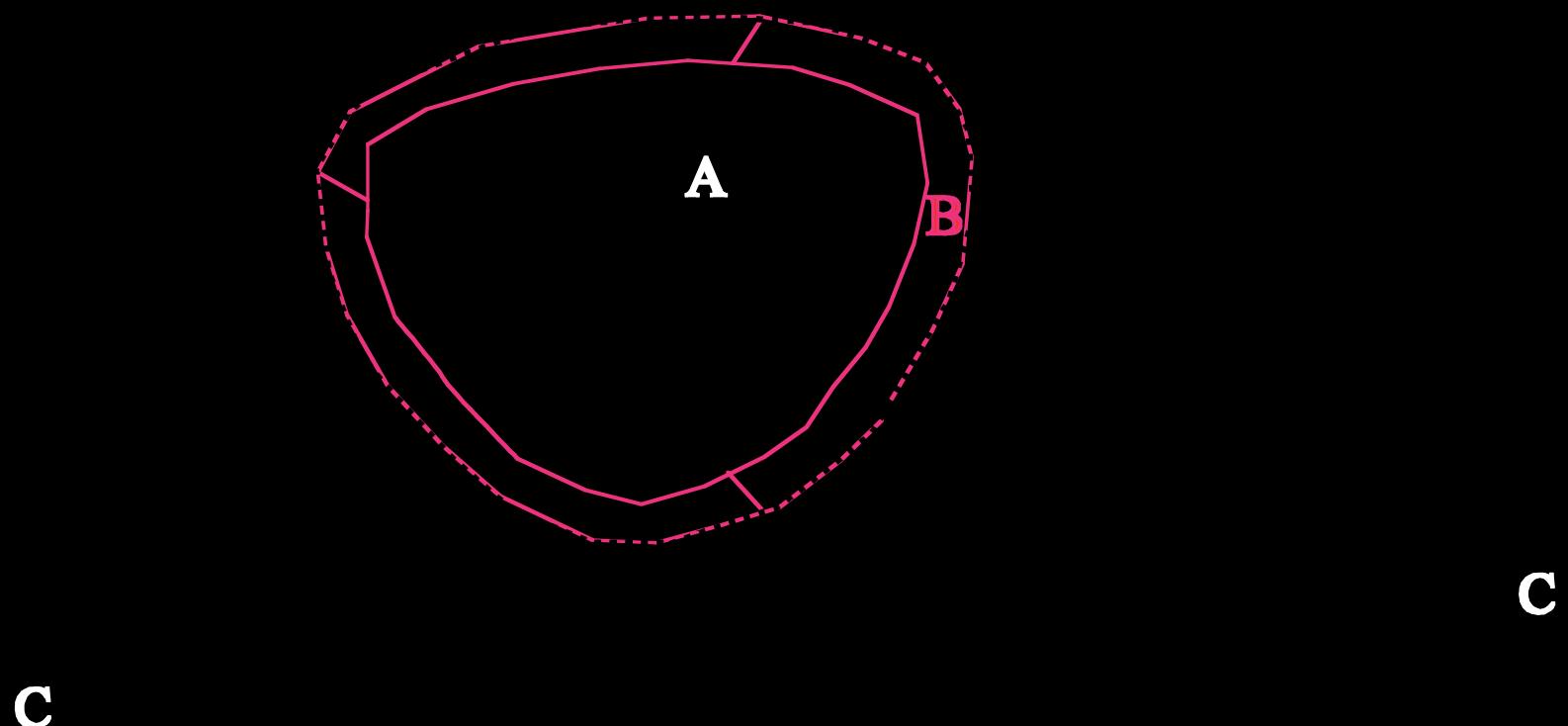


Stage	Map units
1	CC, Cc1, Cc2, Ccc, CEch, CEci, CEd, Cehf, Cf, Cld, Cp, Csc, Ec, Ecc, Eld, Elm, Elph, Em, Emd, Emp, Esc, Ic, Ic2, Icc, Icc2, Ics, Id, Ifc, Im, Im1, Im2, Imd, Ip2, Irc, Isc, It
2	Ia, Ial, Iap, Ic1, Icc1, Ich, Ici, Ico, If, Ig, Ih, Ihe, Ihf, Ihp, Iic, Inbl, Infp, Inp, Int, loc, lohi, Io hn, Ioho, Iohs, Ioht, Iom, Iork, Iorm, Ip, Ip1, Iplt, IpNbm, IpNcl, IpNg, IpNI, IpNt
3	Nb, Nbc, Nbh, Nbl, Nbm, Nc, Ncc, Nhb, Nhc, Nj, Np, NpNbm, NpNbr, NpNhf, NpNt, Npnt, Nsc, Nt, Ntp, pbr, plc, plc1, plc2, plc3, plch, plci, plj, pl1, plp, plr, pNb, pNbm, pNbr, pNc, pNcc, pNt



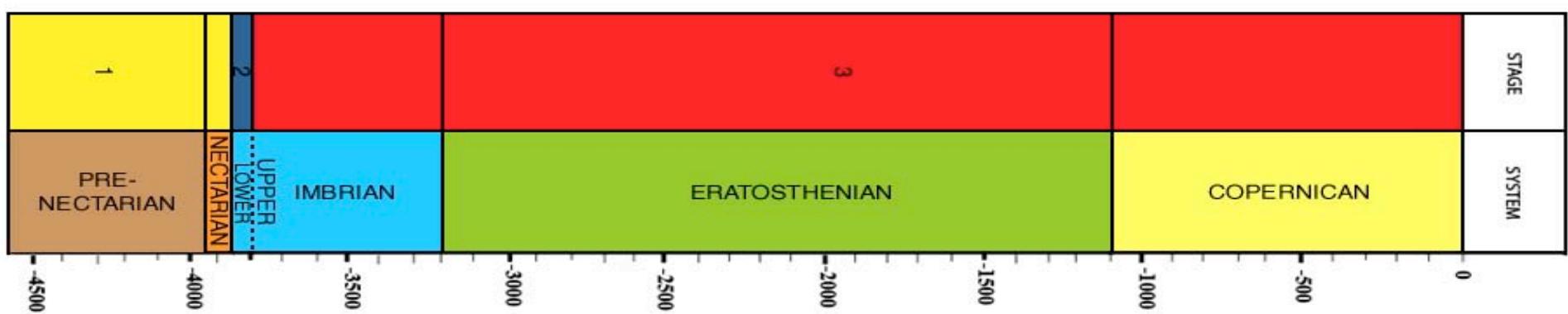
Our approach to investigate surface evolution of the Moon through time

– Apollo-based stratigraphy (3 stages) vs KGRS



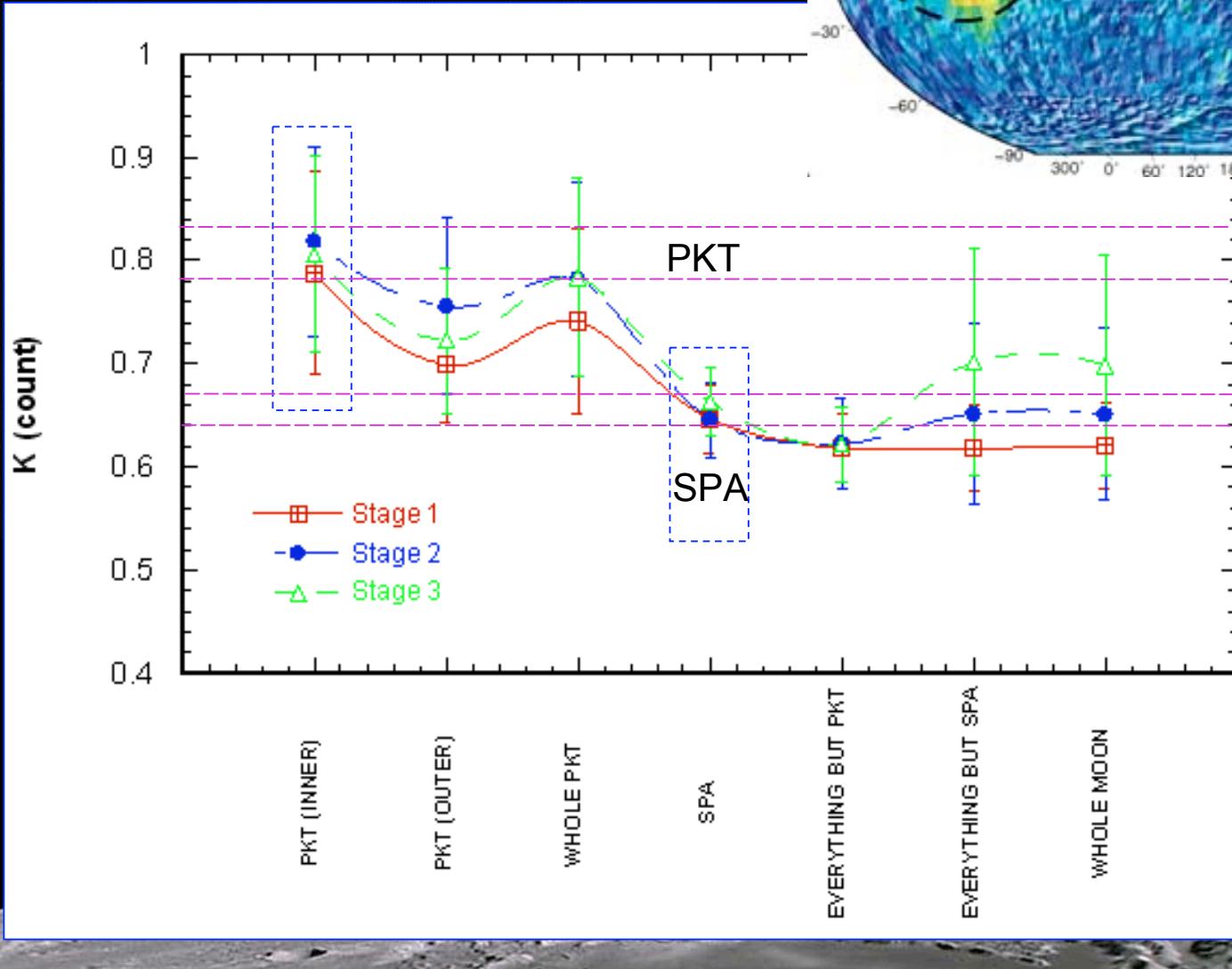
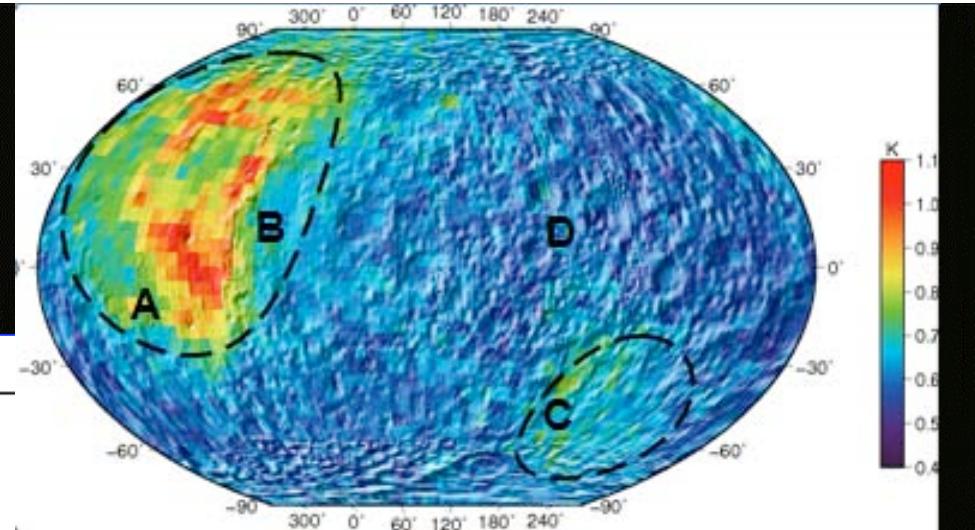


K. J. Kim et al. 2010 (LPSC) and prep.





Comparison of K

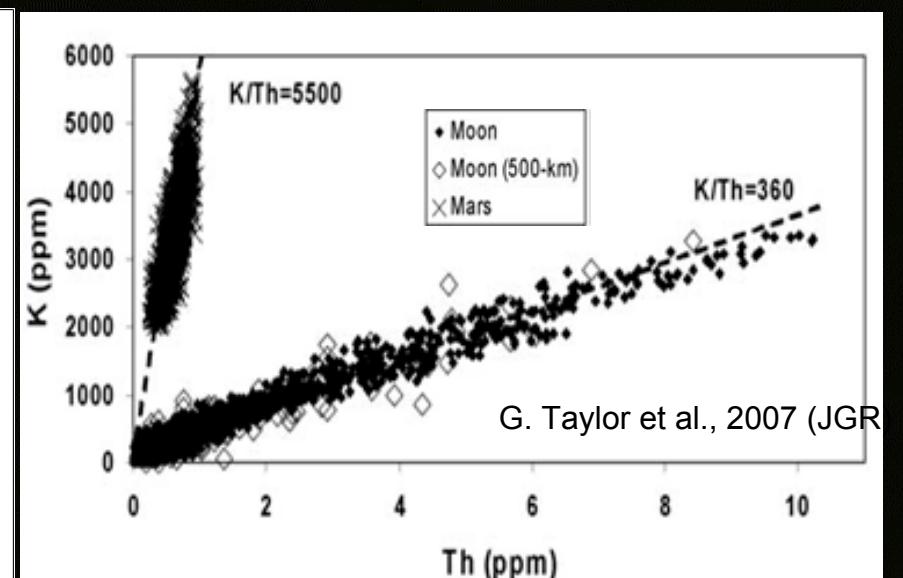
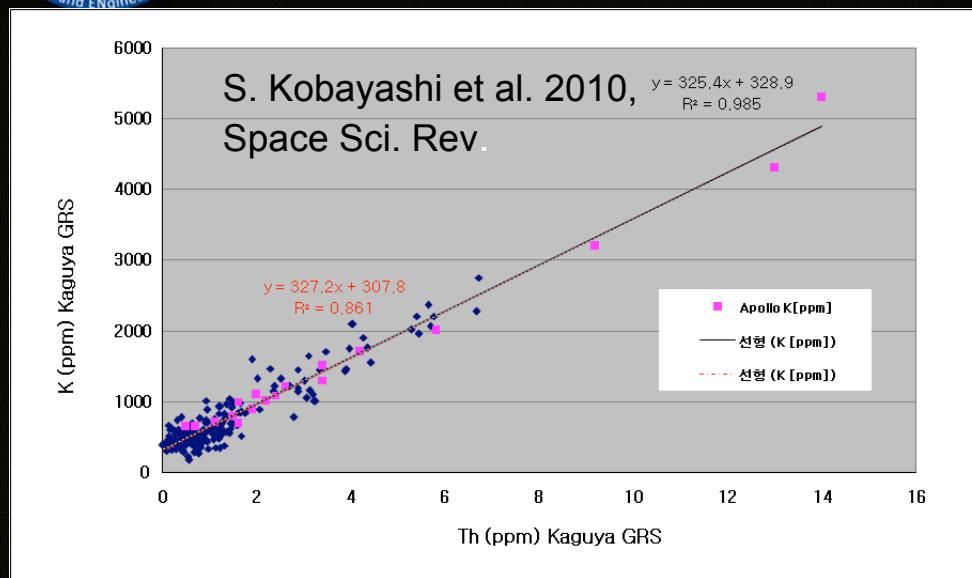


- SPA – K abundance is varied less over time than PKT.
- PKT – Large variation with K over time
→ Continuous activities



Comparison of K and Th

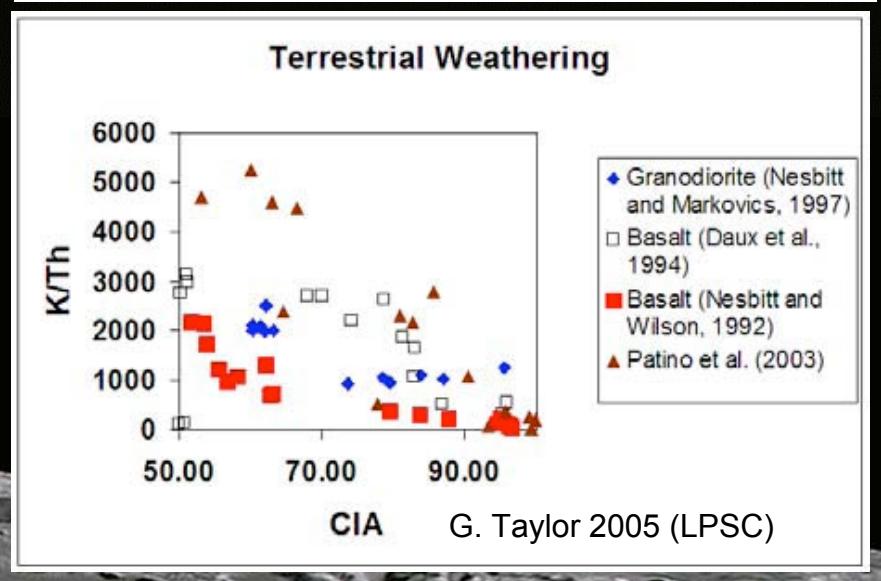
Similar ratio to those in lunar samples and meteorites.



Good correlation among K, Th and U concentrations in the lunar surface :

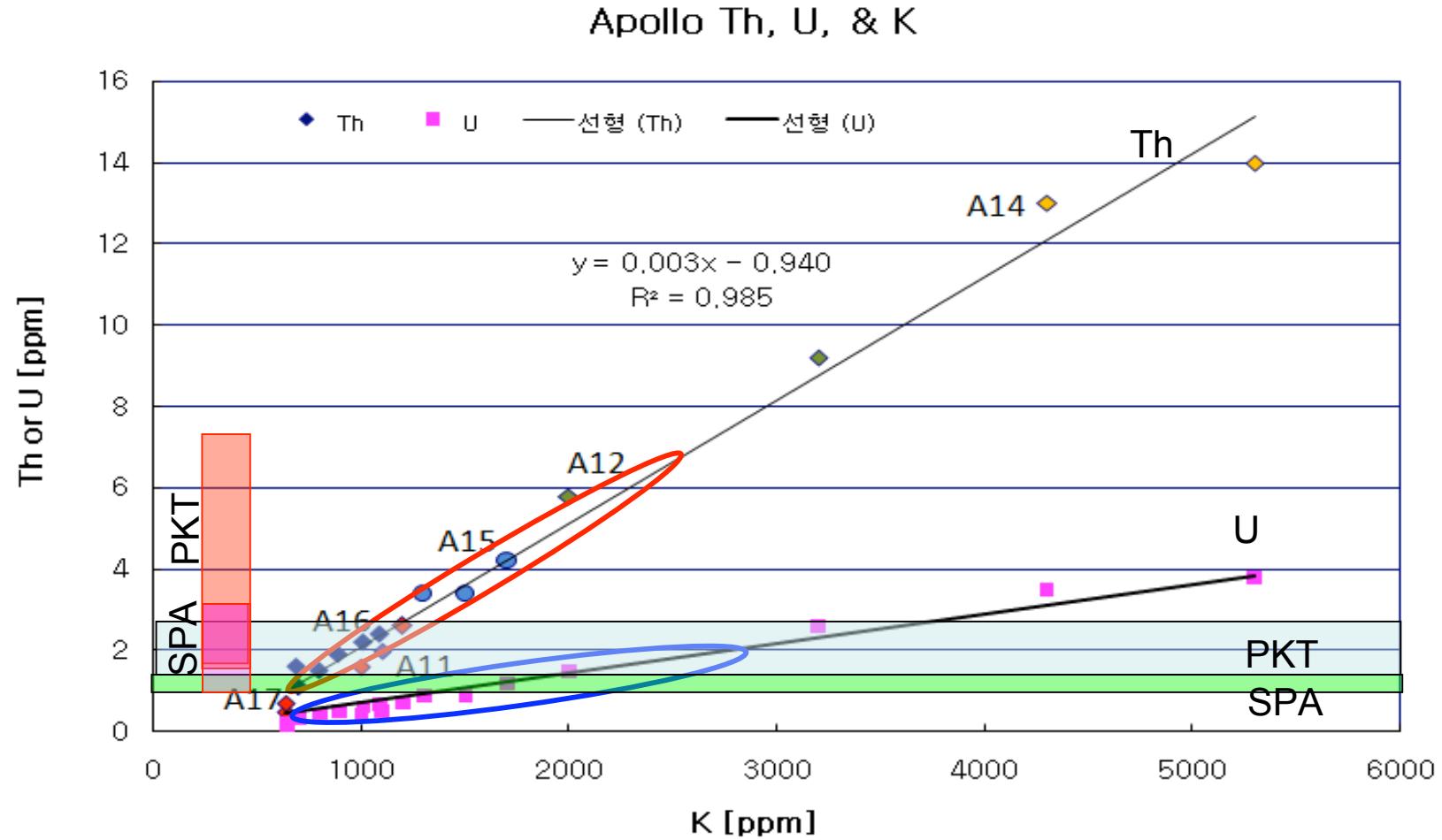
These show global values for K/Th and U/Th. These three elements behave similarly during melting and crystallization in magma. It shows that the Moon is clearly depleted in volatile elements such as K.

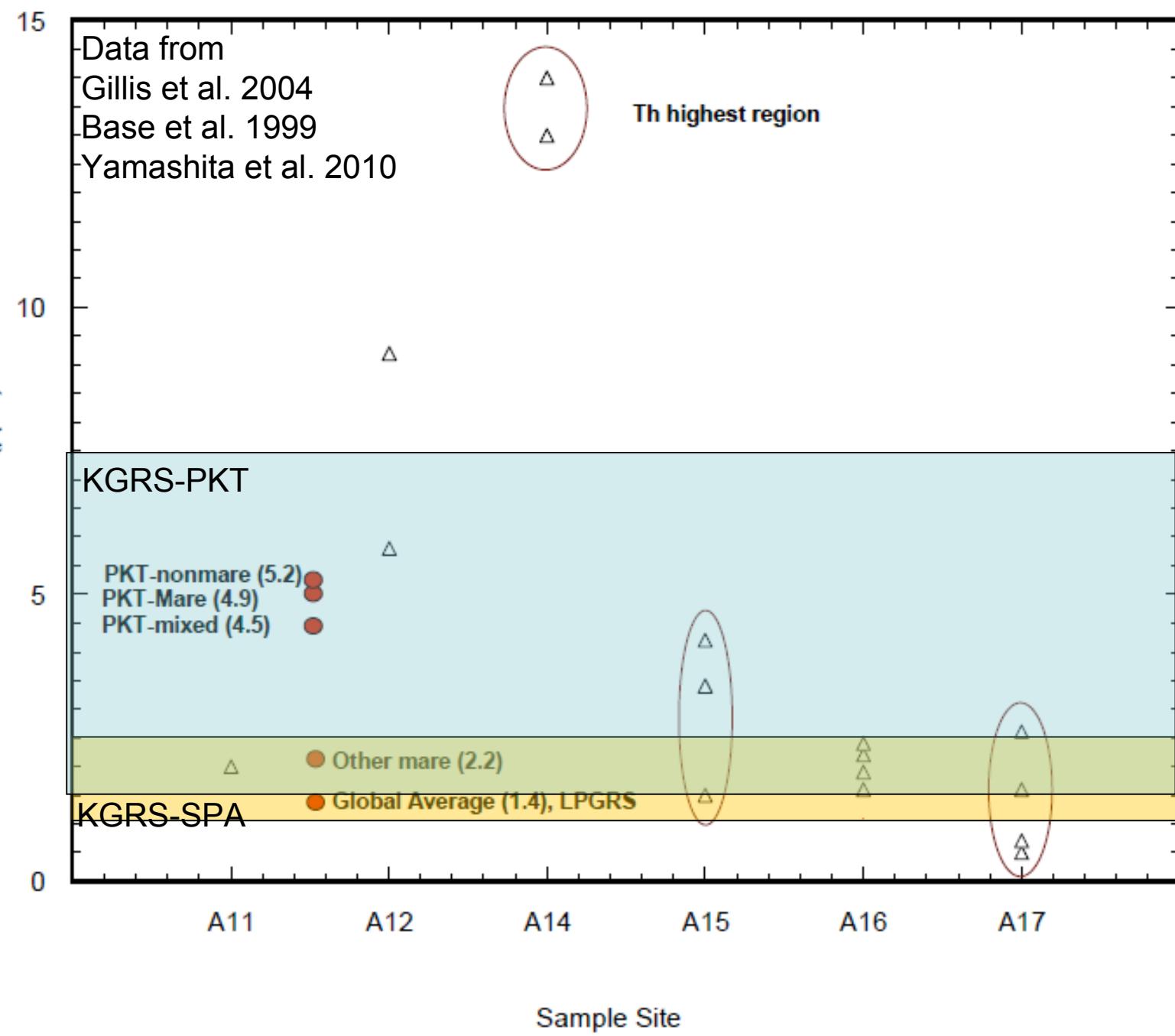
The K/Th ratio is about 360 in the Moon compared to 2900 in Earth and 5300 in Mars.



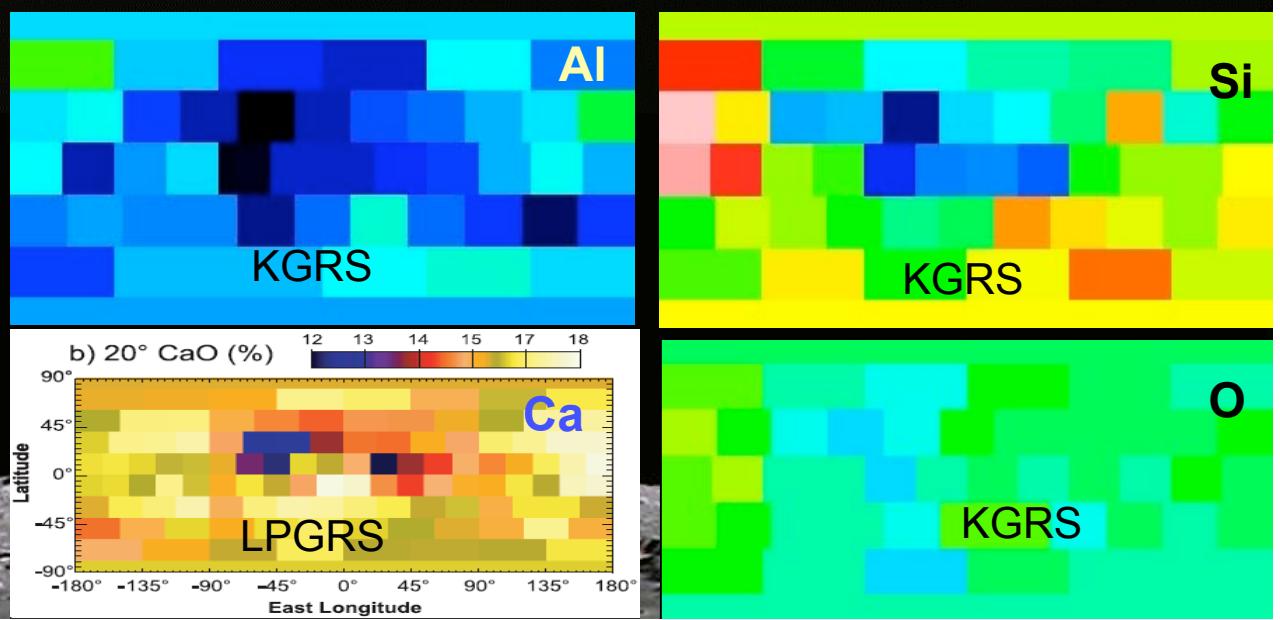
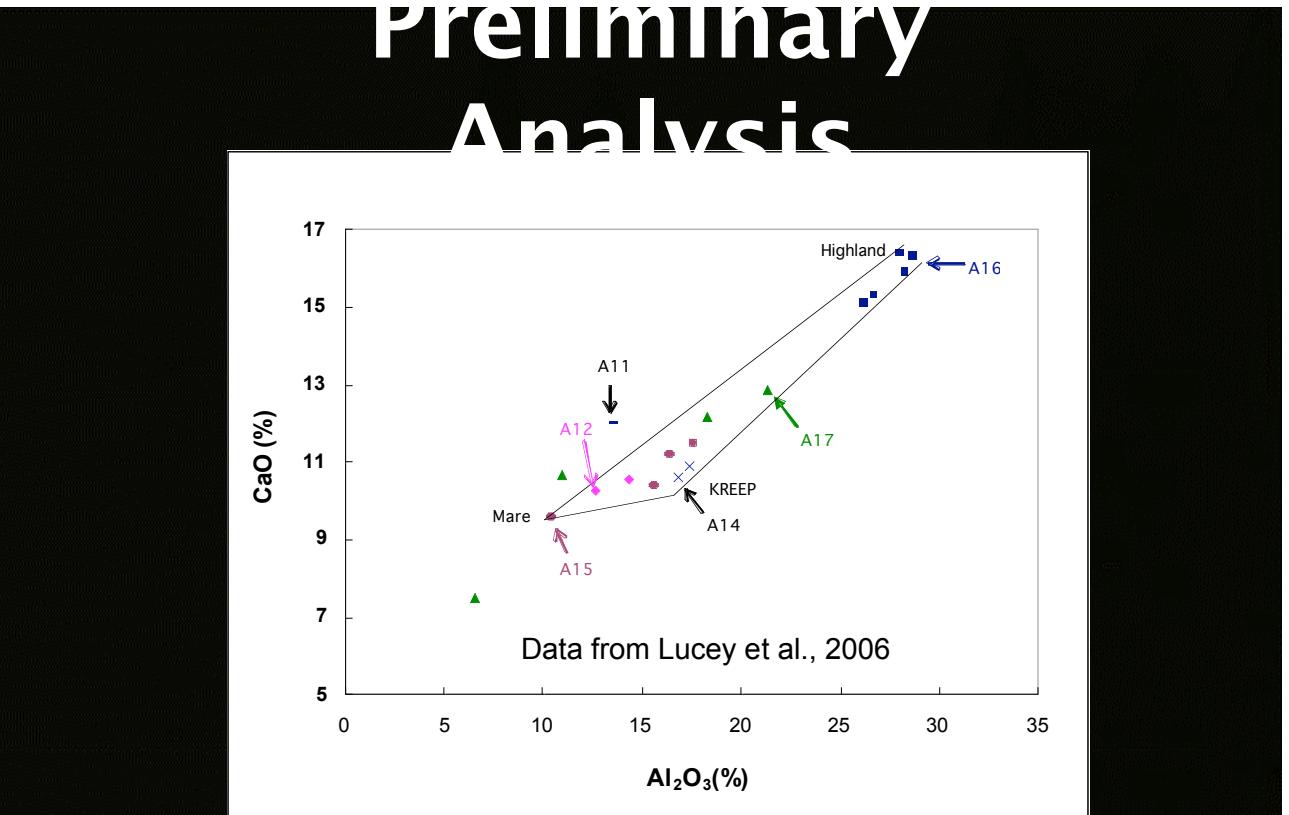
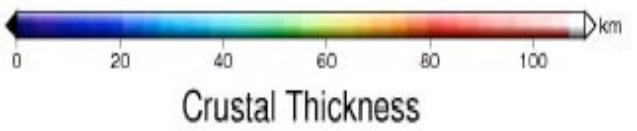
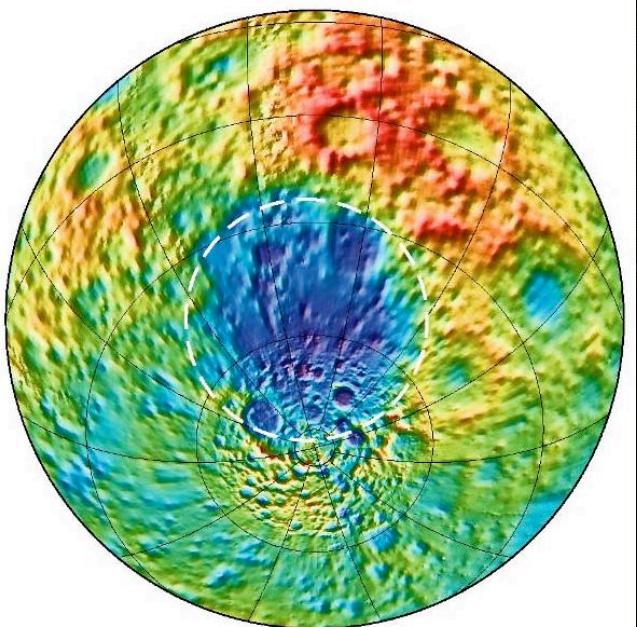
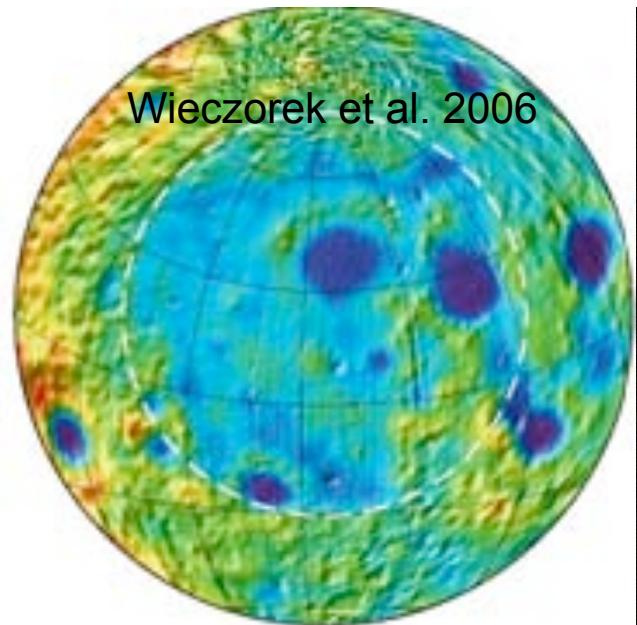


Comparison of K, Th, and U between KGRS and Apollo samples

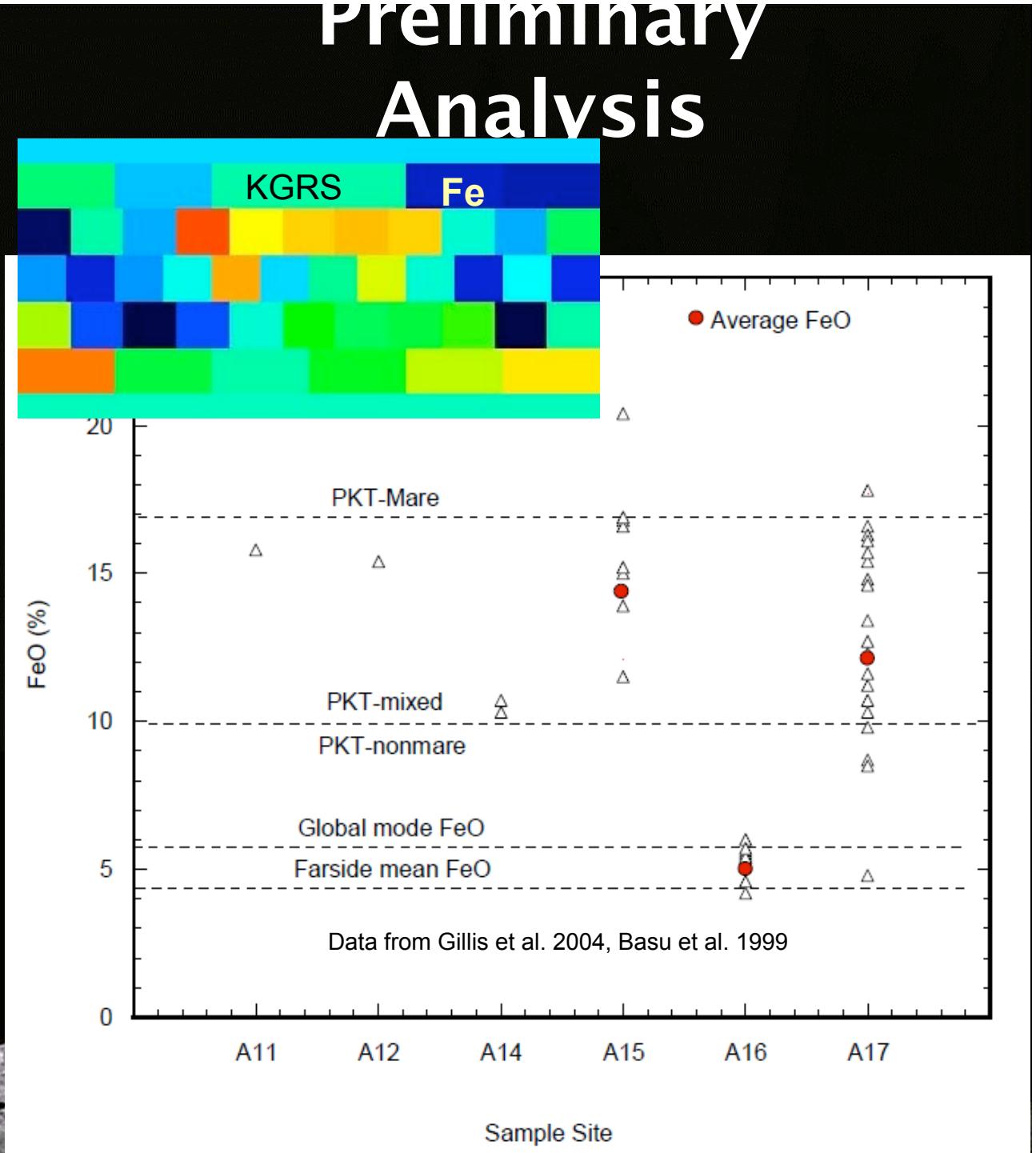
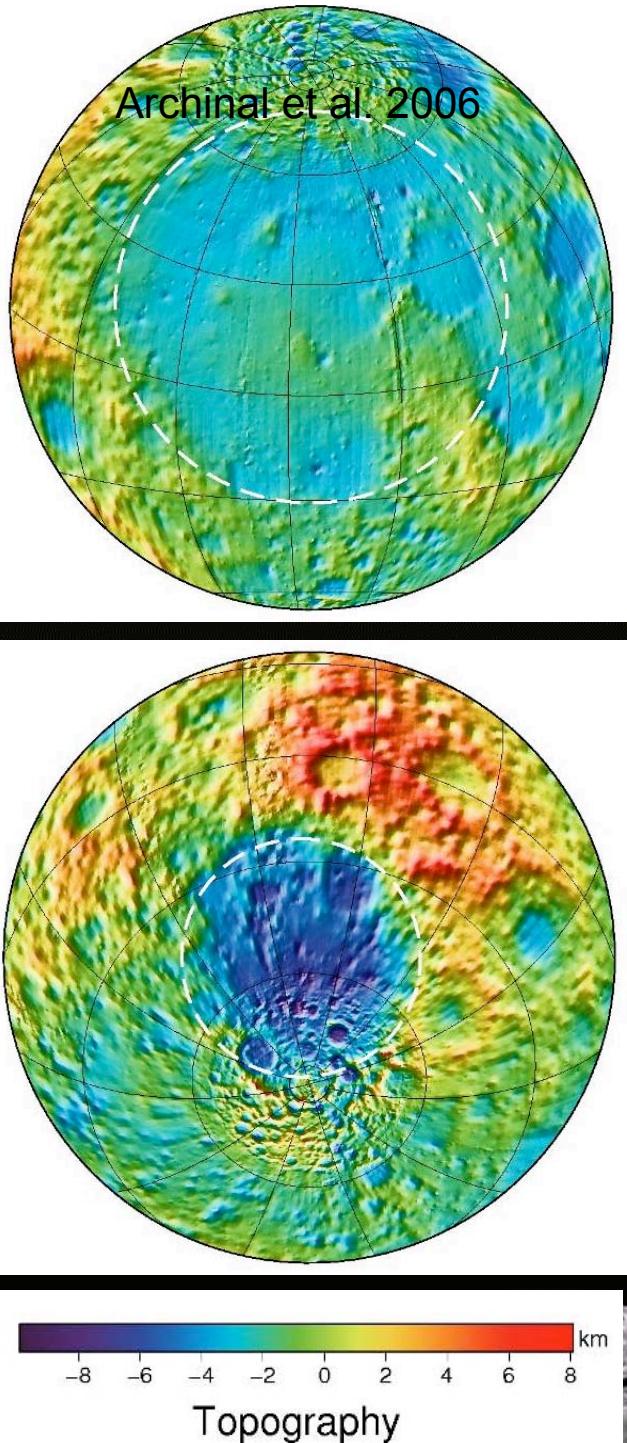




Preliminary Analysis

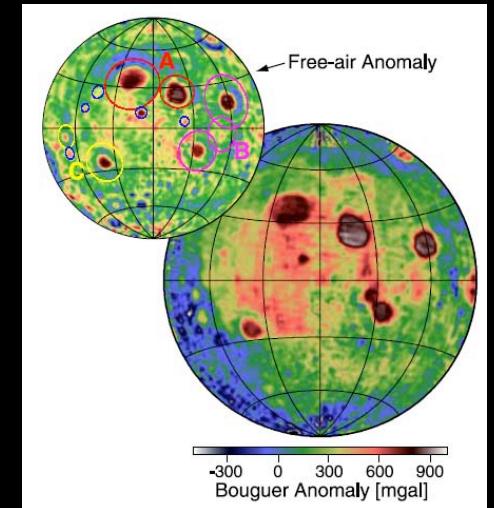


Preliminary Analysis

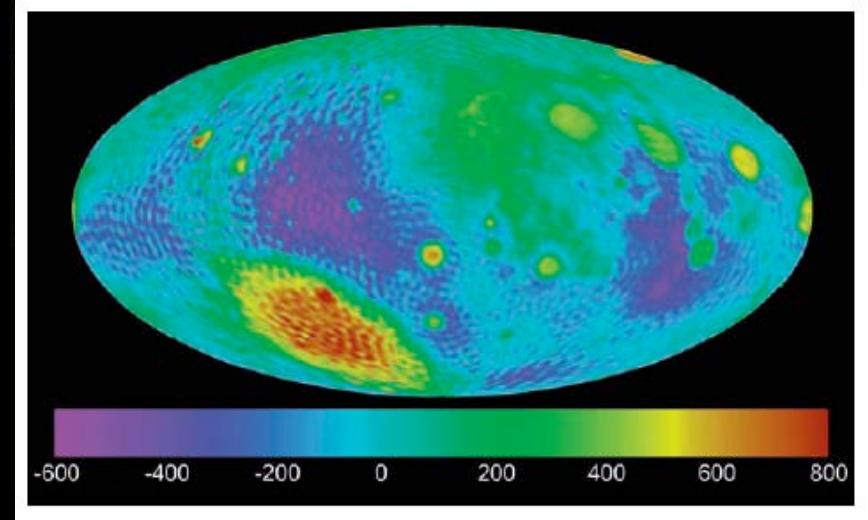
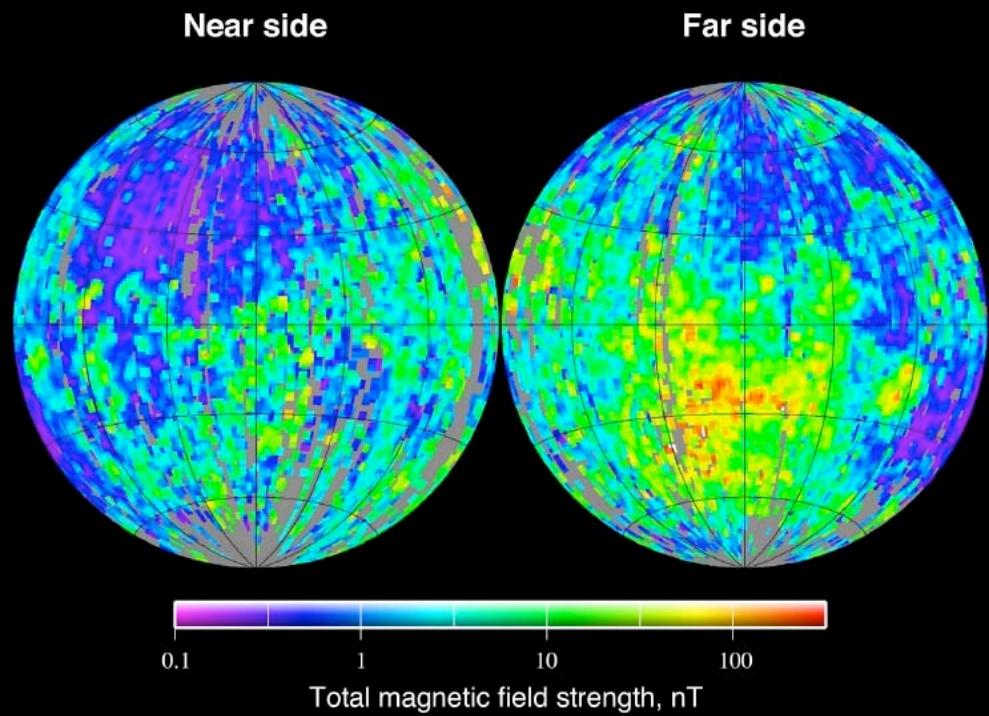


**The South Pole Aitken Basin
has not changed significantly
since the formation.**

**Is the new information of our
study important?**



Sugano & Heki, 2004



Bouguer gravity anomaly map from SGM90d. Units are mGal.

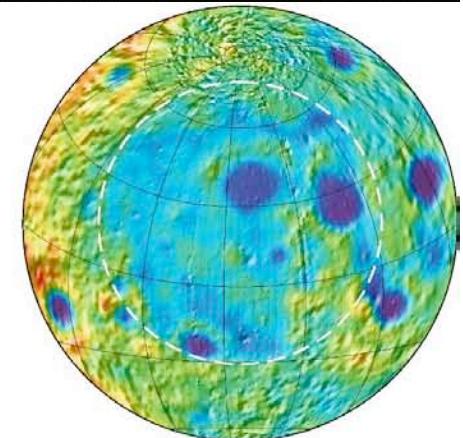
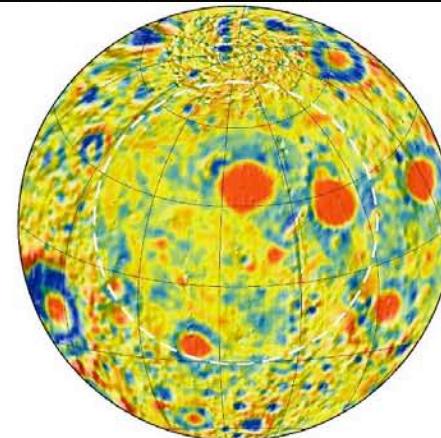
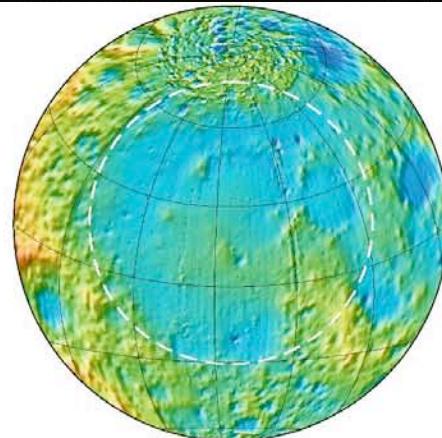
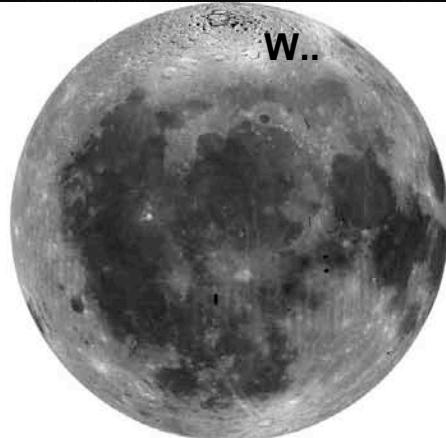
N. Namiki et al., 2010, Science

Lunar Prospector: Electron reflectometer

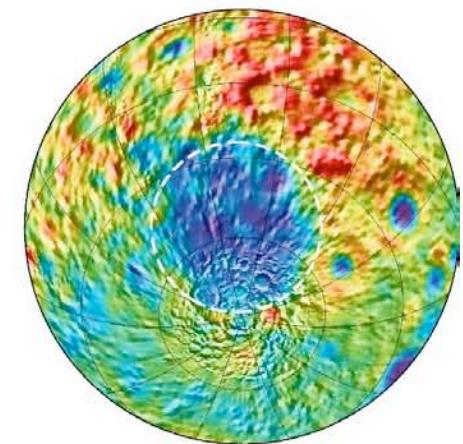
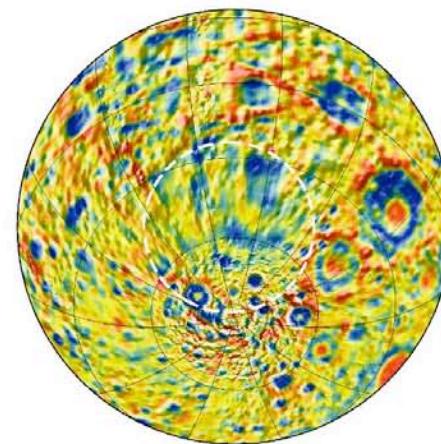
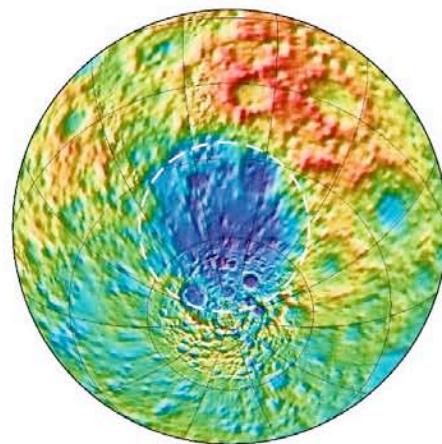


Is Procellarum an Impact Basin [Cadogon (1974) & (Whitaker (1981)] ?

Procellarum Basin



South-Pole Aitken Basin



Clementine UVVIS 750 nm
Wieczorek et al. (2006)

-8 -6 -4 -2 0 2 4 6 8
Topography

-500 -400 -300 -200 -100 0 100 200 300 400 500
Radial Gravity Anomaly

0 20 40 60 80 100
Crustal Thickness

NASA Lunar Science Forum 2010_KJKim



Summary

- Global maps of U, Th and K show SPA has not changed much since its formation similar to other part of the Moon except Procellarum.
- The Procellarum is distinct from the rest of the Moon because of its geologic history (impact?).
- Further investigations of the evolution of lunar surface will continue using additional Kaguya data both elemental and geophysical data.
- Multi perspectives from all existing mission data will be beneficial to pin down the surface evolution of the Moon.



Acknowledgements

We thanks to JAXA and Research Institute of Science and Engineering at Waseda University for providing KGRS data. This work is partially supported by Korea Research Foundation for the Korea-Japan Bilateral Program (09-6303).